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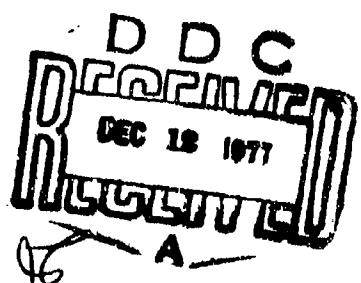
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Algorithm for the Calculation of Absorption Coefficient-Pressure Broadened Molecular Transitions

SHEPARD A. CLOUGH
FRANCIS X. KNEIZYS
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22 July 1977



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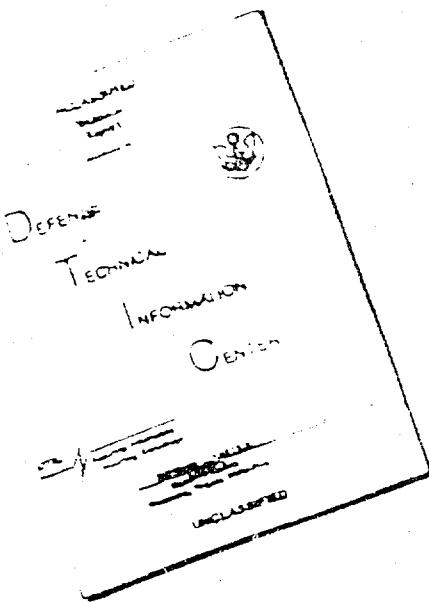
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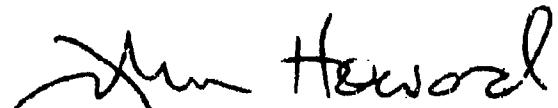


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John Howard
Chief Scientist

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This report describes an algorithm for the accelerated computation of the convolution of a Lorentz line shape (pressure broadened) with spectral line data. A computational savings of approximately 10 has been achieved over conventional methods. The Lorentz function has been decomposed into three functions, each of which is convolved independently at optional sampling intervals. Criteria for the determination of the sampling interval of the Lorentz function for a resultant error of 0.1 percent is described. The report contains		

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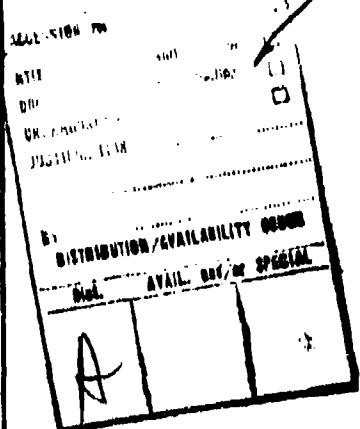
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20. Abstract (Continued)

a listing of the computer program based on the algorithm and sample results
in the spectral region 3550 to 3650 cm⁻¹ due to water and carbon dioxide.



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Algorithm for the Calculation of Absorption Coefficient-Pressure Broadened Molecular Transitions

I. INTRODUCTION

This report describes an algorithm and associated program to perform the convolution of a line shape function with an array of absorption transitions for which the wavenumber values, the intensities, and the halfwidths are specified. The algorithm has been developed with the objectives of minimizing the number of operations to perform the convolution and of developing a program having reasonable computer storage requirements. The decrease in computational operations achieved by the present method may be utilized to decrease computer time requirements for current problems or, if necessary, to perform more extensive calculations that were not previously considered feasible.

The program described in this report uses a given line shape function (the Lorentz function) which is assumed to be independent of molecular species. The algorithm enables the calculation of the absorption coefficient at equally spaced output wavenumber increments over a specified wavenumber region. A uniform path is assumed for which the temperature, pressure, and absorber amounts are specified. The program has been written to be compatible with the AFGL line listing,¹ and performs the function of program LIN previously utilized in HITRAN.
(Received for publication 21 July 1977)

1. McClatchey, R. A., Benedict, W. S., Clough, S. A., Burch, D. E., Caifer, R. F., Fox, K., Rothman, L. S., and Garing, J. S. (1973) AFCRL Atmospheric Absorption Line Parameters Compilation, Report No. AFCRL-TR-73-0096, Environmental Research Papers, No. 434.

calculations. The algorithm has been developed with the goal of keeping the error in the absorption coefficient to within 0.1 percent.

The concept that has been developed involves the decomposition of the given line shape function into three functions which are independently convolved with the parameters for each spectral line. The first function is narrow and describes the central portion of the Lorentzian line shape, the second function is broader describing the intermediate region of the shape, and the third function is broad and may be considered to describe the "near" wing contribution of the line. The functions have been carefully chosen so that the required spectral information is retained as each function is sampled at the same number of points by the convolution process. The spectrum is reconstructed at the final stage by properly interpolating the three functions into the output array of absorption coefficients. Also included in this report is a program that performs a subsequent interpolation providing an essentially continuous resultant spectrum without requiring the convolution calculations to be performed at more output points than are required.

2. DECOMPOSITION OF THE LORENTZ LINE SHAPE

The Lorentz shape is generally applicable for pressure broadened lines for which the transition wavenumber value is large compared to the line halfwidth, and has been discussed by many authors.^{2,3,4} The absorption coefficient, $A_l(\nu)$, as a function of wavenumber, ν , resulting from the convolution of the Lorentz line shape with a molecular transition of wavenumber ν_l , intensity S_l , and width α_l (halfwidth at half maximum: HWHM) is given by

$$A_l(\nu) = \frac{S_l}{\pi} \frac{\alpha_l}{\alpha_l^2 + (\nu - \nu_l)^2} = \frac{S_l}{\pi \alpha_l} \frac{1}{1 + \left(\frac{\nu - \nu_l}{\alpha_l} \right)^2} \quad (1)$$

and the total absorption coefficient is given by

$$A(\nu) = \sum_l A_l(\nu) \quad (2)$$

-
2. Goody, R.M. (1964) Atmospheric Radiation I, Theoretical Basis, Oxford, Clarendon Press.
 3. Barnes, R.B., Jr. (1961) The Shift and Shape and Spectral Lines, Pergamon Press, New York.
 4. Mitchell, A.C.G., and Zemansky, M.W. (1934) Resonance Radiation and Excited Atoms, The Macmillan Co., New York.

For computational purposes, it becomes advantageous to define a dimensionless variable z which for a single line is defined as

$$z = \frac{\nu - \nu_1}{\alpha_1} \quad (3)$$

where z is the wavenumber difference from the line center, ν_1 , in terms of the halfwidth of the transition. The absorption coefficient for a single line in terms of z is

$$A(z) = \frac{S}{\sigma} \frac{1/\pi}{1+z^2} + \left(\frac{S}{\sigma}\right) L(z) \quad (4)$$

where the Lorentz function is

$$L(z) = \frac{1/\pi}{1+z^2} \quad (5)$$

This definition has the computational advantage that $L(z)$ is explicitly independent of halfwidth and may be tabulated as a numerical function with argument z .

The function $L(z)$ has pathologically slow convergence to zero as has been discussed by Goody² and others. In order to avoid computation of the Lorentz function at values of the argument for which the function is slowly varying, the function has been decomposed into three bounded regions ($0 \leq |z| \leq Z_1$, $l = 1, 3$). We have chosen $Z_1 = 3$, $Z_2 = 12$, and $Z_3 = 48$ resulting in three functions with similar functional behavior over the bounds 0 to 3 halfwidths, 0 to 12 halfwidths and 0 to 48 halfwidths respectively.

The decomposition has been performed in a progressive manner by creating the first function (denoted as the "fast function," $(XF(z))$) as the difference between the Lorentz function and an even quartic function $Q_1(z)$ chosen such that the value, the first derivative and the second derivative of $XF(z)$ are zero at the boundary Z_1 . $XF(z)$ is given by

$$XF(z) = 1/\pi (L'(z) - Q_1(z)) \quad \text{for } 0 \leq |z| \leq Z_1 \quad (6)$$

where

$$L'(z) = \pi L(z) = \frac{1}{1+z^2} \quad (7)$$

$$Q_1(z) = a_1 + b_1 z^2 + c_1 z^4 \quad (8)$$

and

$$c_1 = 1/(1+z_b^2)^3, \quad b_1 = -c_1(1+3z_b^2) \quad \text{and} \quad a_1 = c_1(1+3z_b^2+3z_b^4). \quad (9)$$

For $z_b = Z_1 = 3$ we obtain

$$c_1 = \frac{1}{10^3}, \quad b_1 = -\frac{28}{10^3}, \quad \text{and} \quad a_1 = \frac{271}{10^3}. \quad (10)$$

Consequently, we have for the "fast function"

$$XF(z) = 1/\pi (1/(1+z^2) - (a_1 + b_1 z^2 + c_1 z^4)) \quad 0 \leq |z| \leq 3, \quad (11)$$

$$XF(z) = 0 \quad 3 \leq |z|. \quad (12)$$

The second function, the "slow function," $XS(z)$, is constructed in a similar manner by defining an even quartic function, $Q_2(z)$, in the domain $3 \leq |z| \leq 12$ which matches the Lorentzian at $z = Z_b = Z_2 = 12$. We obtain by performing operations similar to those above

$$XS(z) = 1/\pi (L'(z) - Q_2(z)) \quad \text{for} \quad 3 \leq |z| \leq 12 \quad (13)$$

with the coefficients for $Q_2(z)$ given by

$$c_2 = \frac{1}{145^3}, \quad b_2 = -433 \cdot c_2, \quad \text{and} \quad a_2 = 62641 \cdot c_2. \quad (14)$$

The "slow function" is then defined as

$$XS(z) = 1/\pi (Q_1(z) - Q_2(z)) \quad 0 \leq |z| \leq 3, \quad (15)$$

$$XS(z) = 1/\pi (1/(1+z^2) - (a_2 + b_2 z^2 + c_2 z^4)) \quad 3 \leq |z| \leq 12, \quad (16)$$

$$XS(z) = 0 \quad 12 \leq |z|. \quad (17)$$

The third function, the "very slow function," $XVS(z)$, is defined as the Lorentzian for $12 \leq |z| \leq 48$ and as $Q_2(z)$ for $0 \leq |z| \leq 12$.

The following tabular form of the decomposition into the respective domains will make the process more clear:

<u>FUNCTION</u>	<u>DOMAIN</u>		
	$0 \leq z \leq 3$	$3 \leq z \leq 12$	$12 \leq z \leq 48$
$XF(z)$	$\frac{1}{\pi} (L'(z) - Q_1(z))$	0	0
$XS(z)$	$\frac{1}{\pi} (Q_1(z) - Q_2(z))$	$\frac{1}{\pi} (L'(z) - Q_2(z))$	0
$XVS(z)$	$\frac{1}{\pi} Q_2(z)$	$\frac{1}{\pi} Q_2(z)$	$\frac{1}{\pi} L'(z)$

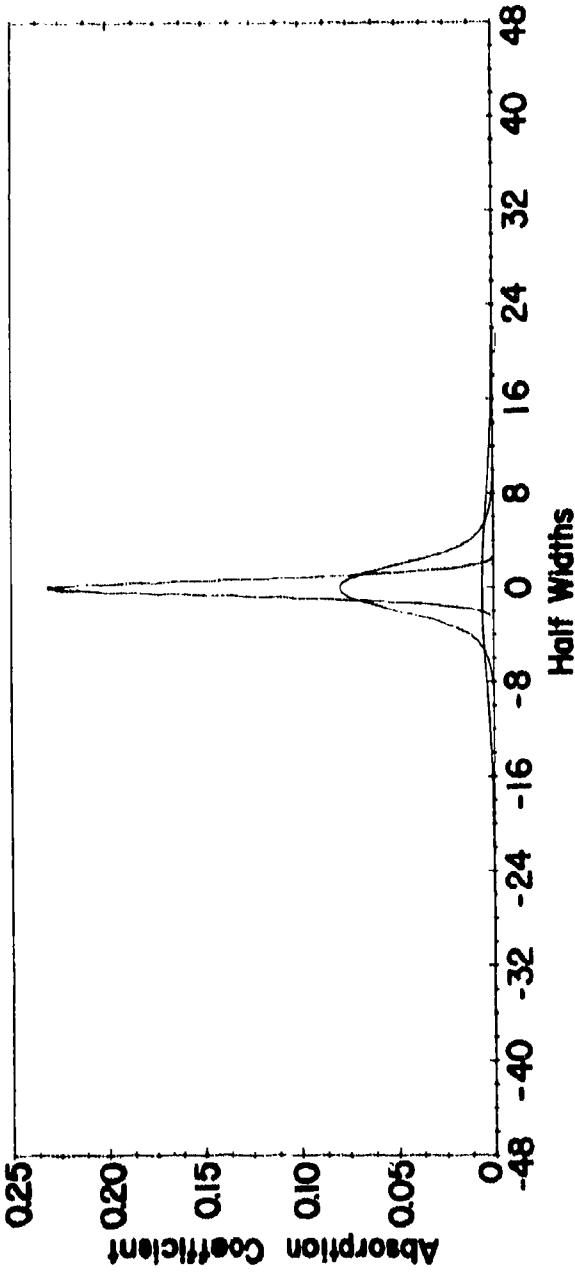
(18)

Note that the three functions sum to the Lorentzian in each domain and that the functions are continuous in value, first and second derivative across the boundaries of the domains.

While the choices of the functions and the boundaries may appear to be somewhat arbitrary, an examination of the result demonstrates the advantage of the imposed conditions. Figure 1 is a plot of the three functions from -48 to +48 half-widths and indicates the dependence of $XF(z)$, $XS(z)$, and $XVS(z)$ as a function of the variable z . It should be noted that 40.71 percent of the integral of the Lorentz function is contained in XF , 42.47 percent in XS , and 15.50 percent in XVS . The remaining 1.32 percent is outside the 48 halfwidth cutoff.

The three functions are represented in Figure 2 with the horizontal axis determined by the boundary of each function and the vertical axis chosen appropriately for each function. The three functions in this representation are seen to be functionally similar and as a consequence, it is possible to sample each function at the same sampling interval to perform the convolution. One hundred and fifty one (151) values of each symmetric function are stored in the computer program, and it is these values that are plotted in Figure 2 as indicated by the index on the top horizontal axis. The tabulation of this number of values for the functions precludes the necessity for interpolation in the convolution calculation. The values of the functions are tabulated in Table 1.

LORENTZ DECOMPOSITION



10

Figure 1. Decomposition of the Lorentz Function into Three Domains: -3 to 3 halfwidths, -12 to 12 halfwidths and -48 to 48 halfwidths. The sum of the three functions gives $L(z) = (1/\pi)(1/(1 + z^2))$

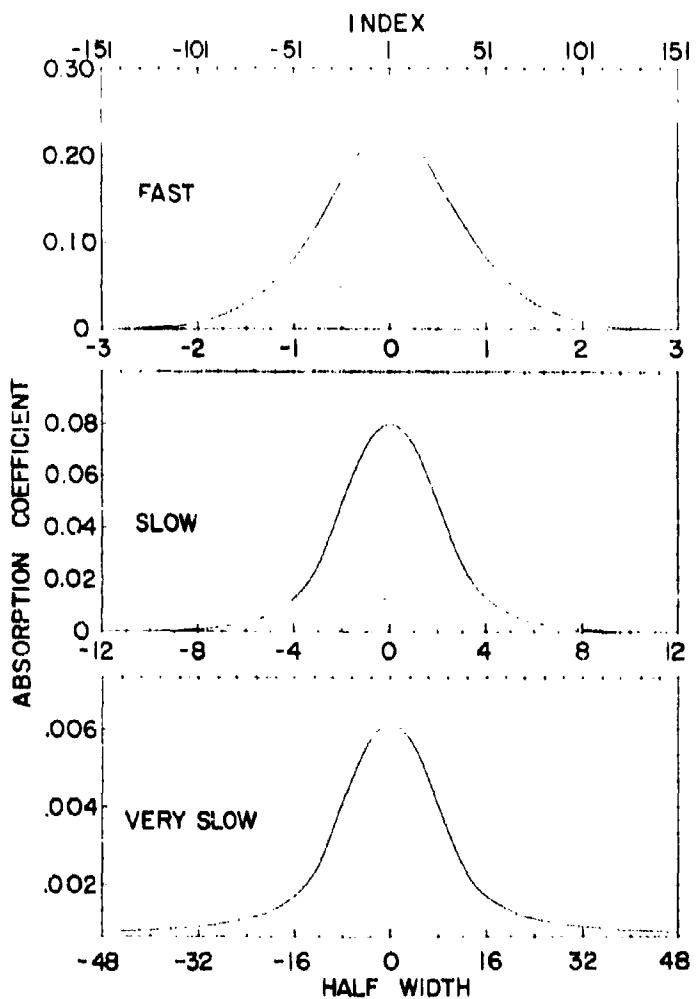


Figure 2. Functional Behavior of the Decomposed Functions Over Their Respective Limits. Plots are direct plots of the functions in terms of the storage index of the program. The fast function (XF) is on the top, the slow function (XS) is in the middle, and the very slow functions (XVS) are on the bottom

Table 1. Tabulation of the Decomposed Functions Used to Describe the Lorentz Function. The variable z is defined as the wavenumber difference from the line center in terms of the halfwidth, α , that is, $z = |(\nu - \nu_0)/\alpha|$. Plots of the three tabulated functions appear in Figure 2 and Figure 3.

INDEX	"FAST" FUNCTION		"SLOW" FUNCTION		"VERY SLOW" FUNCTION	
	z	FN	z	FN	z	FN
1	0.030	2.320E-01	0.000	7.972E-02	3.000	6.540E-03
2	.020	2.319E-01	.050	7.966E-02	.320	6.536E-03
3	.040	2.316E-01	.150	7.949E-02	.640	6.522E-03
4	.060	2.309E-01	.240	7.921E-02	.960	6.499E-03
5	.080	2.301E-01	.320	7.882E-02	1.280	6.467E-03
6	.100	2.290E-01	.420	7.811E-02	1.600	6.425E-03
7	.120	2.277E-01	.520	7.770E-02	1.920	6.375E-03
8	.140	2.261E-01	.560	7.697E-02	2.240	6.316E-03
9	.160	2.243E-01	.640	7.614E-02	2.560	6.249E-03
10	.180	2.223E-01	.720	7.521E-02	2.880	6.173E-03
11	.200	2.202E-01	.800	7.418E-02	3.200	6.088E-03
12	.220	2.178E-01	.880	7.305E-02	3.520	5.996E-03
13	.240	2.152E-01	.950	7.182E-02	3.840	5.896E-03
14	.260	2.125E-01	1.040	7.050E-02	4.160	5.789E-03
15	.280	2.096E-01	1.120	6.910E-02	4.480	5.675E-03
16	.300	2.066E-01	1.200	6.781E-02	4.800	5.554E-03
17	.320	2.034E-01	1.280	6.659E-02	5.120	5.427E-03
18	.340	2.001E-01	1.360	6.441E-02	5.440	5.294E-03
19	.360	1.967E-01	1.440	6.270E-02	5.760	5.155E-03
20	.380	1.932E-01	1.520	6.093E-02	6.080	5.012E-03
21	.400	1.896E-01	1.600	5.911E-02	6.400	4.864E-03
22	.420	1.859E-01	1.680	5.723E-02	6.720	4.712E-03
23	.440	1.821E-01	1.760	5.531E-02	7.040	4.556E-03
24	.460	1.783E-01	1.840	5.335E-02	7.360	4.398E-03
25	.480	1.745E-01	1.920	5.136E-02	7.680	4.237E-03
26	.500	1.706E-01	2.000	4.934E-02	8.000	4.075E-03
27	.520	1.667E-01	2.080	4.731E-02	8.320	3.911E-03
28	.540	1.628E-01	2.160	4.528E-02	8.640	3.747E-03
29	.560	1.588E-01	2.240	4.324E-02	8.960	3.584E-03
30	.580	1.549E-01	2.320	4.121E-02	9.280	3.421E-03
31	.600	1.510E-01	2.400	3.920E-02	9.600	3.261E-03
32	.620	1.470E-01	2.480	3.722E-02	9.920	3.103E-03
33	.640	1.432E-01	2.560	3.527E-02	10.240	2.948E-03
34	.660	1.393E-01	2.640	3.330E-02	10.560	2.797E-03
35	.680	1.355E-01	2.720	3.133E-02	10.880	2.652E-03
36	.700	1.317E-01	2.800	2.976E-02	11.200	2.512E-03
37	.720	1.279E-01	2.880	2.806E-02	11.520	2.379E-03
38	.740	1.242E-01	2.960	2.646E-02	11.840	2.255E-03
39	.760	1.205E-01	3.040	2.495E-02	12.160	2.138E-03
40	.780	1.169E-01	3.120	2.354E-02	12.480	2.031E-03
41	.800	1.134E-01	3.200	2.223E-02	12.800	1.931E-03
42	.820	1.099E-01	3.280	2.100E-02	13.120	1.839E-03
43	.840	1.065E-01	3.360	1.966E-02	13.440	1.752E-03
44	.860	1.031E-01	3.440	1.878E-02	13.760	1.672E-03
45	.880	9.984E-02	3.520	1.779E-02	14.080	1.598E-03
46	.900	9.661E-02	3.600	1.683E-02	14.400	1.528E-03
47	.920	9.345E-02	3.680	1.594E-02	14.720	1.462E-03
48	.940	9.035E-02	3.760	1.511E-02	15.040	1.401E-03
49	.960	8.733E-02	3.840	1.432E-02	15.360	1.343E-03
50	.980	8.437E-02	3.920	1.359E-02	15.680	1.289E-03
51	1.000	8.149E-02	4.000	1.286E-02	16.000	1.239E-03
52	1.020	7.867E-02	4.080	1.222E-02	16.320	1.191E-03
53	1.040	7.592E-02	4.160	1.160E-02	16.640	1.145E-03
54	1.060	7.324E-02	4.240	1.101E-02	16.960	1.103E-03
55	1.080	7.063E-02	4.320	1.046E-02	17.280	1.062E-03
56	1.100	6.809E-02	4.400	9.930E-03	17.600	1.024E-03
57	1.120	6.561E-02	4.480	9.432E-03	17.920	9.882E-04
58	1.140	6.320E-02	4.560	8.960E-03	18.240	9.539E-04
59	1.160	6.086E-02	4.640	8.513E-03	18.560	9.214E-04
60	1.180	5.858E-02	4.720	8.089E-03	18.880	8.909E-04

Table 1. Tabulation of the Decomposed Functions Used to Describe the Lorentz Function. The variable z is defined as the wavenumber difference from the line center in terms of the halfwidth, α , that is, $z = [(v - v_0)/\alpha]$. Plots of the three tabulated functions appear in Figure 2 and Figure 3 (Cont.)

INDEX	"FAST" FUNCTION		"SLOW" FUNCTION		"VERY SLOW" FUNCTION	
	Z	FN	Z	FN	Z	FN
61	1.200	5.637E-02	4.800	7.687E-03	19.200	8.611E-04
62	1.220	5.422E-02	4.880	7.305E-03	19.520	8.332E-04
63	1.240	5.213E-02	4.960	6.942E-03	19.840	8.066E-04
64	1.260	5.010E-02	5.040	6.597E-03	20.160	7.813E-04
65	1.280	4.813E-02	5.120	6.269E-03	20.480	7.571E-04
66	1.300	4.622E-02	5.200	5.958E-03	20.800	7.344E-04
67	1.320	4.437E-02	5.280	5.661E-03	21.120	7.120E-04
68	1.340	4.250E-02	5.360	5.379E-03	21.440	6.910E-04
69	1.360	4.064E-02	5.440	5.111E-03	21.760	6.708E-04
70	1.380	3.915E-02	5.520	4.855E-03	22.080	6.515E-04
71	1.400	3.752E-02	5.600	4.611E-03	22.400	6.331E-04
72	1.420	3.594E-02	5.680	4.379E-03	22.720	6.155E-04
73	1.440	3.441E-02	5.760	4.158E-03	23.040	5.985E-04
74	1.460	3.293E-02	5.840	3.947E-03	23.360	5.822E-04
75	1.480	3.150E-02	5.920	3.746E-03	23.680	5.666E-04
76	1.500	3.012E-02	6.000	3.555E-03	24.000	5.517E-04
77	1.520	2.879E-02	6.080	3.372E-03	24.320	5.373E-04
78	1.540	2.749E-02	6.160	3.198E-03	24.640	5.234E-04
79	1.560	2.625E-02	6.240	3.032E-03	24.960	5.101E-04
80	1.580	2.504E-02	6.320	2.873E-03	25.280	4.973E-04
81	1.600	2.388E-02	6.400	2.722E-03	25.600	4.850E-04
82	1.620	2.276E-02	6.480	2.578E-03	25.920	4.731E-04
83	1.640	2.168E-02	6.560	2.441E-03	26.240	4.616E-04
84	1.660	2.064E-02	6.640	2.309E-03	26.560	4.506E-04
85	1.680	1.963E-02	6.720	2.184E-03	26.880	4.399E-04
86	1.700	1.865E-02	6.800	2.065E-03	27.200	4.297E-04
87	1.720	1.773E-02	6.880	1.951E-03	27.520	4.197E-04
88	1.740	1.684E-02	6.960	1.843E-03	27.840	4.102E-04
89	1.760	1.597E-02	7.040	1.739E-03	28.160	4.009E-04
90	1.780	1.514E-02	7.120	1.641E-03	28.480	3.920E-04
91	1.800	1.435E-02	7.200	1.547E-03	28.800	3.833E-04
92	1.820	1.358E-02	7.280	1.457E-03	29.120	3.749E-04
93	1.840	1.284E-02	7.360	1.372E-03	29.440	3.668E-04
94	1.860	1.214E-02	7.440	1.291E-03	29.760	3.591E-04
95	1.880	1.146E-02	7.520	1.213E-03	30.080	3.514E-04
96	1.900	1.081E-02	7.600	1.140E-03	30.400	3.441E-04
97	1.920	1.019E-02	7.680	1.070E-03	30.720	3.369E-04
98	1.940	9.594E-03	7.760	1.003E-03	31.040	3.303E-04
99	1.960	9.024E-03	7.840	9.397E-04	31.360	3.233E-04
100	1.980	8.479E-03	7.920	8.796E-04	31.680	3.164E-04
101	2.000	7.958E-03	8.000	8.224E-04	32.000	3.105E-04
102	2.020	7.460E-03	8.080	7.682E-04	32.320	3.044E-04
103	2.040	6.985E-03	8.160	7.167E-04	32.640	2.985E-04
104	2.060	6.532E-03	8.240	6.679E-04	32.960	2.927E-04
105	2.080	6.101E-03	8.320	6.217E-04	33.280	2.871E-04
106	2.100	5.690E-03	8.400	5.779E-04	33.600	2.817E-04
107	2.120	5.299E-03	8.480	5.365E-04	33.920	2.764E-04
108	2.140	4.928E-03	8.560	4.973E-04	34.240	2.713E-04
109	2.160	4.575E-03	8.640	4.603E-04	34.560	2.663E-04
110	2.180	4.241E-03	8.720	4.254E-04	34.880	2.614E-04
111	2.200	3.924E-03	8.800	3.925E-04	35.200	2.567E-04
112	2.220	3.624E-03	8.880	3.615E-04	35.520	2.521E-04
113	2.240	3.341E-03	8.960	3.323E-04	35.840	2.476E-04
114	2.260	3.073E-03	9.040	3.049E-04	36.160	2.433E-04
115	2.280	2.821E-03	9.120	2.793E-04	36.480	2.390E-04
116	2.300	2.584E-03	9.200	2.550E-04	36.800	2.349E-04
117	2.320	2.361E-03	9.280	2.324E-04	37.120	2.308E-04
118	2.340	2.152E-03	9.360	2.113E-04	37.440	2.269E-04
119	2.360	1.958E-03	9.440	1.910E-04	37.760	2.231E-04
120	2.380	1.773E-03	9.520	1.732E-04	38.080	2.194E-04

Table 1. Tabulation of the Decomposed Functions Used to Describe the Lorentz Function. The variable z is defined as the wavenumber difference from the line center in terms of the halfwidth, α , that is, $z = [(\nu - \nu_0)/\alpha]$. Plots of the three tabulated functions appear in Figure 2 and Figure 3 (Cont.)

INDEX	"FAST" FUNCTION		"SLOW" FUNCTION		"VERY SLOW" FUNCTION	
	z	FN	z	FN	z	FN
121	2.400	1.602E-03	9.600	1.561E-04	38.400	2.157E-04
122	2.420	1.642E-03	9.680	1.403E-04	39.720	2.122E-04
123	2.440	1.294E-03	9.760	1.256E-04	39.040	2.087E-04
124	2.460	1.157E-03	9.840	1.121E-04	39.360	2.053E-04
125	2.480	1.030E-03	9.920	9.955E-05	39.680	2.020E-04
126	2.500	9.131E-04	10.000	8.806E-05	40.000	1.988E-04
127	2.520	8.055E-04	10.080	7.753E-05	40.320	1.957E-04
128	2.540	7.070E-04	10.160	6.791E-05	40.640	1.926E-04
129	2.560	6.170E-04	10.240	5.919E-05	40.960	1.896E-04
130	2.580	5.352E-04	10.320	5.121E-05	41.280	1.867E-04
131	2.600	4.610E-04	10.400	4.403E-05	41.600	1.838E-04
132	2.620	3.942E-04	10.480	3.758E-05	41.920	1.811E-04
133	2.640	3.343E-04	10.560	3.182E-05	42.240	1.783E-04
134	2.660	2.809E-04	10.640	2.669E-05	42.560	1.756E-04
135	2.680	2.336E-04	10.720	2.215E-05	42.880	1.730E-04
136	2.700	1.920E-04	10.800	1.810E-05	43.200	1.705E-04
137	2.720	1.557E-04	10.880	1.472E-05	43.520	1.680E-04
138	2.740	1.244E-04	10.960	1.174E-05	43.840	1.655E-04
139	2.760	9.758E-05	11.040	9.194E-06	44.160	1.631E-04
140	2.780	7.498E-05	11.120	7.054E-06	44.480	1.608E-04
141	2.800	5.620E-05	11.200	5.280E-06	44.800	1.585E-04
142	2.820	4.088E-05	11.280	3.834E-06	45.120	1.563E-04
143	2.840	2.869E-05	11.360	2.693E-06	45.440	1.541E-04
144	2.860	1.915E-05	11.440	1.791E-06	45.760	1.519E-04
145	2.880	1.203E-05	11.520	1.124E-06	46.080	1.498E-04
146	2.900	6.947E-06	11.600	6.479E-07	46.400	1.478E-04
147	2.920	3.549E-06	11.680	3.306E-07	46.720	1.458E-04
148	2.940	1.494E-05	11.760	1.790E-07	47.040	1.438E-04
149	2.960	4.618E-07	11.840	4.104E-08	47.360	1.419E-04
150	2.980	5.511E-08	11.920	5.113E-09	47.680	1.400E-04
151	3.000	0.	12.000	0.	48.000	1.381E-04

The rationalization for terminating the convolution at 48 halfwidths is based on three considerations: (1) evidence that the line shape deviates from Lorentz behavior for $(\nu - \nu_0)$ of the order of 5 cm^{-1} , corresponding to $z = 48$ for a typical atmospheric line at atmospheric pressure (Winters et al.;⁵ Holstein;⁶ Breene;³ Burch;⁷) (2) the value of the Lorentzian at 48 halfwidths is 1.38×10^{-4} which is

5. Winters, B. H., Silverman, S., and Benedict, W. S. (1964) Line shape in the wing beyond the band head of the 4.3μ band of CO_2 , J. Quant. Spect. Rad. Trans., 4:527.
6. Holstein, T. (1950) Pressure broadening of spectral lines, Phys. Rev. 79:744; see also L. Spitzer Phys. Rev. 58:348 (1940).
7. Burch, D. E., Grynak, D. A., Fatty, R. R. and Bartky, C. E. (1968) The Shapes of Collision Broadened CO_2 Lines, Philco Ford Corp., Aeronutronic Report U-3203.

sufficiently small for most calculations, and (3) calculation beyond 48 halfwidths requires an unwarranted increase in computational effort. Although the value of the Lorentz function may justify termination of the calculation at 48 halfwidths, it must not be forgotten that 1.32 percent of the integrated absorption is beyond this region. This is a result of the previously mentioned slow convergence of $1/z^2$ (Goody,² and Townes and Schawlow⁶). It is suggested that calculations of the far wings of the lines can more appropriately be done in a parametrized tabulation of what is currently referred to as the continuum for each of the molecular species of interest. The contribution of this continuum function to the absorption coefficient may be calculated at large sampling intervals in the wavenumber domain with negligible effect on computational time.

3. SAMPLING INTERVAL.

General discussions of the proper choice of the sampling interval appear in the literature for the type of "line by line" calculations described in this report.^{9, 10, 11, 12} These references are primarily concerned with the effect of the sampling interval on the integral of the absorption coefficient or absorptance over a specified wavenumber interval. That approach is particularly appropriate to the case in which an instrumental scanning function that is broad compared to the spectral line widths, is being convolved with the true line spectrum. The present discussion is concerned with the development of a quantitative evaluation of the error introduced in the reconstruction of the spectrum in the interval between sampled points.

An approach that provides direct insight into the problem of determining the sampling interval is that obtained from information theory and the Nyquist sampling theorem. Simple and readily comprehensible discussions of this are provided by

8. Townes, C. H., and Schawlow, A. L. (1955) Microwave Spectroscopy, McGraw Hill Book Co., New York.
9. Drayson, D. R. (1967) The Calculation of Long Wave Radiative Transfer in Planetary Atmospheres, Report No. 07584-1-T, University of Michigan, Ann Arbor.
10. Kyle, T. G. (1968) Net interval for calculating absorption spectra, J. Opt. Soc. Am. 58:192.
11. Scott, N. A. (1974) A direct method of computation of the transmission function of an inhomogeneous gaseous medium, I: Description of the method, J. Quant. Spectr. Rad. Trans. 14:691.
12. Kunde, V. G., and McGuire, W. C. (1974) Direct integration transmittance model, J. Quant. Spectr. Rad. Trans. 14:803.

Brigham¹³ and Bracewell.¹⁴ The sampling theorem states that if the Fourier transform of a function is zero for all frequencies greater than frequency f_o , then the continuous function $g(x)$ can be uniquely determined from its sampled values, $g_n(n\delta x)$, if the sampling interval is chosen as

$$\delta x = \frac{1}{2f_o} . \quad (19)$$

In particular, the continuous function $g(x)$ can be reconstructed for all x by convolving the continuous interpolating function

$$s(x) = \frac{1}{\delta x} \frac{\sin(\pi \frac{x}{\delta x})}{\left(\pi \frac{x}{\delta x}\right)} \quad (20)$$

with the sampled function $g_n(n\delta x)$ where n is integer. That is,

$$g(x) = g_n(n\delta x) * s(x) \quad (21)$$

or explicitly for the discrete convolution we have

$$g(x) = \sum_{n=-\infty}^{\infty} g_n(n\delta x) \frac{\sin \pi \left(\frac{x}{\delta x} - n \right)}{\pi \left(\frac{x}{\delta x} - n \right)} . \quad (22)$$

In the present case, the Fourier transform of the Lorentz function does not have a frequency, f_o , beyond which the transform is zero. Consequently by sampling at discrete intervals, it is not possible to exactly reconstruct the continuous function. However, by properly choosing the interval, the difference between the exact and reconstructed functions can be held within a predetermined limit. The Lorentz function is given as

$$g(x) = \frac{1}{\pi} \frac{\alpha}{\alpha^2 + x^2} \quad (23)$$

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- 13. Brigham, E.O. (1974) The Fast Fourier Transform, Englewood Cliffs, New Jersey.
 - 14. Bracewell, R.M. (1986) The Fourier Transform and Its Applications, McGraw-Hill, New York.

where $x = \nu - \nu_0$, α is the halfwidth, and ν_0 is the wavenumber value of the transition. The Fourier transform of $g(x)$ is

$$\mathcal{F}(g(x)) = g(f) = e^{-2\pi\alpha f} \quad (24)$$

where \mathcal{F} is the Fourier transform operator. If the value of the transform is assumed to be zero beyond a frequency, f_0 , this is equivalent to multiplying the complete transform by a rectangular box, $\mathcal{L}(f)$, where

$$\begin{aligned} 1 & \quad |f| < f_0 \\ \mathcal{L}(f) &= \frac{1}{2} \quad |f| = f_0 \\ 0 & \quad |f| > f_0 \end{aligned} \quad . \quad (25)$$

We may define the sampling interval as before such that

$$\delta x = \frac{1}{2f_0} \quad . \quad (26)$$

The Fourier transform of $\mathcal{L}(f)$ in the x domain is given by

$$s(x) = \frac{1}{\delta x} \frac{\sin(\pi \frac{x}{\delta x})}{(\pi \frac{x}{\delta x})} \quad (27)$$

which is identical to the scanning function prescribed by the Nyquist sampling theorem (Eq. (20)).

In order to evaluate the error caused by the discrete sampling interval and the associated truncation of the function in the Fourier domain, calculations have been made using the interpolating function, Eq. (27), convolved with the sampled Lorentz function

$$g_s(n\delta x) = \frac{1}{\pi} \frac{\alpha}{\alpha^2 + (n\delta x)^2} \quad (28)$$

as indicated in Eq. (22). For these convolution calculations, a more relevant definition of the sampling interval, δx , is given in terms of the halfwidth, α ,

$$\delta x = \frac{\alpha}{N} \quad . \quad (29)$$

The results are evaluated for sampling intervals expressed in fractions, $1/N$, of the halfwidth. For this definition, we have for the sampled Lorentz function

$$g_N \left(\alpha \frac{n}{N} \right) = \frac{1}{\pi \alpha} \frac{1}{1 + \left(\frac{n}{N} \right)^2} \quad (30)$$

A quantitative evaluation of the effect of the sampling interval, is the RMS deviation of the function interpolated at three intermediate points per interval from the function calculated exactly at the same intermediate points. The results of these calculations are given in Table 2. These results are a measure of the amount of information lost due to the choice of sampling interval. As the sampling interval becomes infinitesimally small, the RMS deviation approaches zero. It is seen that for an interval of $\alpha/4$, that is, $N = 4$, the RMS deviation is negligibly small and essentially all of the information is retained.

Table 2. Effect of Sampling Interval on the Reconstruction of the Lorentz Function Using $(\sin x/x)$ Interpolation and Four Point Lagrange Interpolation. Interpolations are calculated at three equally spaced points between the sampled values. In all cases the maximum percent deviation using the four point Lagrange occurs in the first sampled interval. Deviations are calculated as the difference between the interpolated value and the exact value of the function

Sampling Interval		Interpolation Method		
		$\frac{\sin x}{x}$	Four Point Lagrange	
N	$\delta x = \frac{\alpha}{N}$	rms deviation	rms deviation	max percent deviation
1	α	1.1×10^{-3}	3.5×10^{-3}	-1.9
2	$\frac{\alpha}{2}$	0.067	0.20	-0.99
3	$\frac{\alpha}{3}$	0.0035	0.087	-0.37
4	$\frac{\alpha}{4}$	0.00063	0.026	-0.15

There is another consideration that must be taken into account in this situation. Although the information is retained for a sampling interval of $\alpha/4$, the computational effort to retrieve this information using the $(\sin x/x)$ interpolation is large and in some cases may be of the same magnitude as that required to evaluate the function directly at the intermediate intervals.

In order to retrieve sufficient information with reasonable computational effort and stay within a specified error criterion for the absorption coefficient calculation, a four point interpolation function has been used instead of the prohibitively long "sin x/x" interpolation. The output points are assumed to be equally spaced. Values for the weighting factors necessary to perform the interpolation at three equally spaced intermediate points are given by the following three equations:

$$g\left(x + \frac{\delta x}{4}\right) = -\frac{7}{128} g(x - \delta x) + \frac{105}{128} g(x) + \frac{35}{128} g(x + \delta x) - \frac{5}{128} g(x + 2\delta x) \quad (31)$$

$$g\left(x + \frac{\delta x}{2}\right) = -\frac{1}{16} g(x - \delta x) + \frac{9}{16} g(x) + \frac{1}{16} g(x + \delta x) - \frac{1}{16} g(x + 2\delta x) \quad (32)$$

$$g\left(x + \frac{3\delta x}{4}\right) = \frac{5}{28} g(x - \delta x) + \frac{35}{128} g(x) + \frac{105}{128} g(x + \delta x) - \frac{7}{128} g(x + 2\delta x) \quad (33)$$

The interpolation scheme that has been employed in this development is due to Lagrange and is well described by Acton.¹⁵ In Table 2, the maximum percent deviation in the interpolated values is indicated. This error always occurs within one or two intervals from the line center. For a sampling interval of $\alpha/4$, the error is 0.15 percent. Although this error at the line center is slightly larger than desirable, the added complication of using a more refined interpolation does not presently seem warranted. Furthermore, the percent error decreases rapidly away from the line center.

The four point interpolation scheme is used for interpolating the "very slow array," and the "slow array" into the "fast array." This is required since the sampling criterion of $\alpha/4$ has been used for the calculation of the three arrays where the effective α is appropriate to each array.

Similarly, for the presentation of the final results, the fast array is interpolated into a plotting array in a separate program which represents the final result. Figure 3 depicts the effect of this interpolation scheme as compared to the result obtained with linear interpolation. It should be noted that the error for a complete spectrum can only be the same or less than that for a single line. This is readily concluded from the superposition theorem.

One final consideration that provides information on the sampling interval, is the Rayleigh criterion.¹⁶ For two spectral lines of unit strength and equal half-widths, separated by two halfwidths, let us determine the sampling interval required to exactly reproduce the minimum between the two spectral lines.

15. Acton, F. S. (1970) Numerical Methods That Work, New York.

16. Born, M., and Wolf, E. (1986) Principles of Optics, Oxford, Pergamon Press.

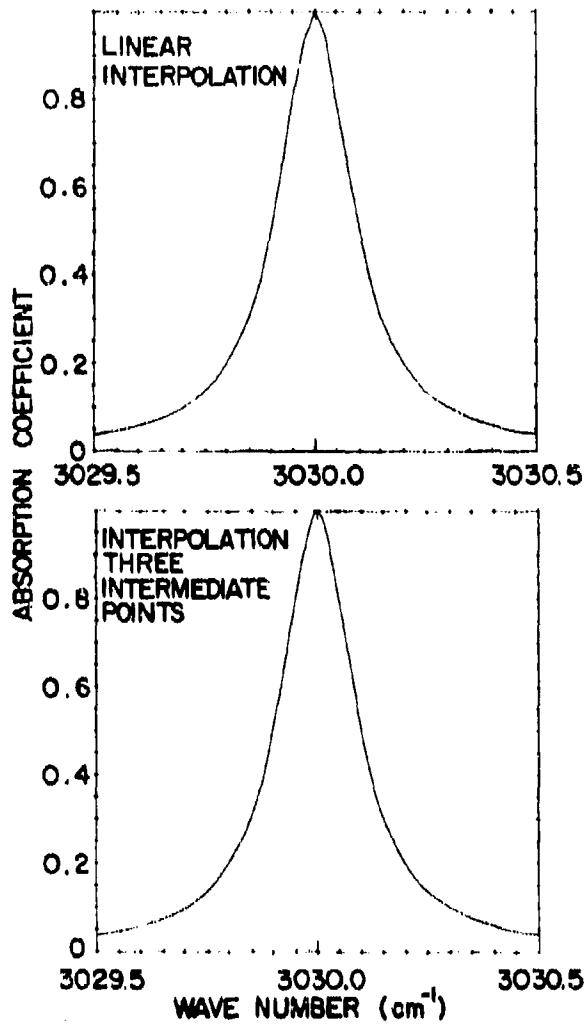


Figure 3. Reconstruction of the Shape of a Single Line of Strength π , halfwidth, $\alpha = 0.1$, Using a Sampling Interval of $\delta x = \alpha/4$ With a Sampling Point Positioned at the Spectral Line Center. The upper curve uses linear interpolation of the convolution output and the lower curve uses the 4 point Lagrange interpolation. To emphasize the line center, the region from -5 to +5 halfwidths is depicted

Consider one line centered at $x = 0$ and the other line centered at $x = x_0$. We have for the two Lorentz functions

$$g_1(x) = \frac{1}{\pi} \frac{\alpha}{\alpha^2 + x^2} \quad (34)$$

and

$$g_2(x) = \frac{1}{\pi} \frac{\alpha}{\alpha^2 + (x - x_0)^2} \quad (35)$$

The Fourier transform for this pair is given by

$$g(f) = e^{-2\pi\alpha f} + e^{-2\pi\alpha f} \cos(2\pi x_0 f) \quad (36)$$

The Rayleigh criterion postulates $x_0 = 2\alpha$ giving for the transform

$$g(f) = e^{-2\pi\alpha f} (1 + \cos 4\pi\alpha f) \quad (37)$$

To exactly reproduce the minimum between the doublet, one period of the oscillatory function is required, giving

$$4\pi\alpha f_0 = 2\pi \quad (38)$$

so that the cut off frequency is

$$f_0 = \frac{1}{2\alpha} \quad (39)$$

and the sampling interval from the Nyquist theorem, is

$$\delta x = \frac{1}{2f_0} = \alpha \quad (40)$$

This result indicates that the sampling interval of $\alpha/4$, previously determined, is well within the Rayleigh criterion for resolving a Lorentzian doublet separated by two halfwidths.

4. APPLICATION OF THE ALGORITHM

In the previous sections, two basic steps for an improved computational procedure to perform the convolution of the Lorentz line shape with a set of absorption line data have been described: (1) the decomposition of the line shape into bounded domains, and (2) the criterion for determining the sampling interval. In this section we describe the application of the algorithm to obtain the absorption coefficient as a function of wavenumber for a set of spectral line data parameters that are appropriate for the specified values of temperature, pressure, and column density. Three separate convolutions utilizing the fast (XF), the slow (XS), and the very slow (XVS) functions are performed on the data set. The results of the three convolutions for a specific set of data are displayed in transmittance in Figure 4. For this case the average halfwidth is of the order of 0.04 cm^{-1} so that the sampling interval used is $\delta x = \alpha/4 = 0.01 \text{ cm}^{-1}$. Each convolution function is sampled at approximately 24 values corresponding to intervals of 0.01 cm^{-1} for the fast function, 0.04 cm^{-1} for the slow function, and 0.16 cm^{-1} for the very slow function. Seventy-two (72) points have been sampled to reproduce the Lorentzian from -48 to +48 halfwidths instead of 384 (4×96) which would be required if the function were sampled at equal intervals of $\alpha/4$ (0.01 cm^{-1}). Consequently, the algorithm as applied in this case provides a computational savings of a factor of 1/3. The actual saving using an optimized version of the conventional method compared to the present method is a factor of 10. The additional factor of two is due to the use of tabulated rather than calculated functions and to reduced overhead time in setting up the inner convolution loop. The results of the two computational methods are given in Figure 5 and no differences are discernable. It should be noted that the lower spectrum of Figure 5 is equivalent to the product of the three functions presented in Figure 4, since the functions are all on a transmittance scale.

Some further consideration needs to be given to the method described here with respect to the theoretical computational gain. If the assumption is made that the line shape function can be decomposed into M functions, and that the sampling interval is the same for the first decomposed function as for the total Lorentz function, $\alpha/4$, then the gain, G, is

$$G = \frac{N_{\text{tot}}}{\sum_m^M N_m \left(\frac{\delta x_1}{\delta x_m} \right)} . \quad (41)$$

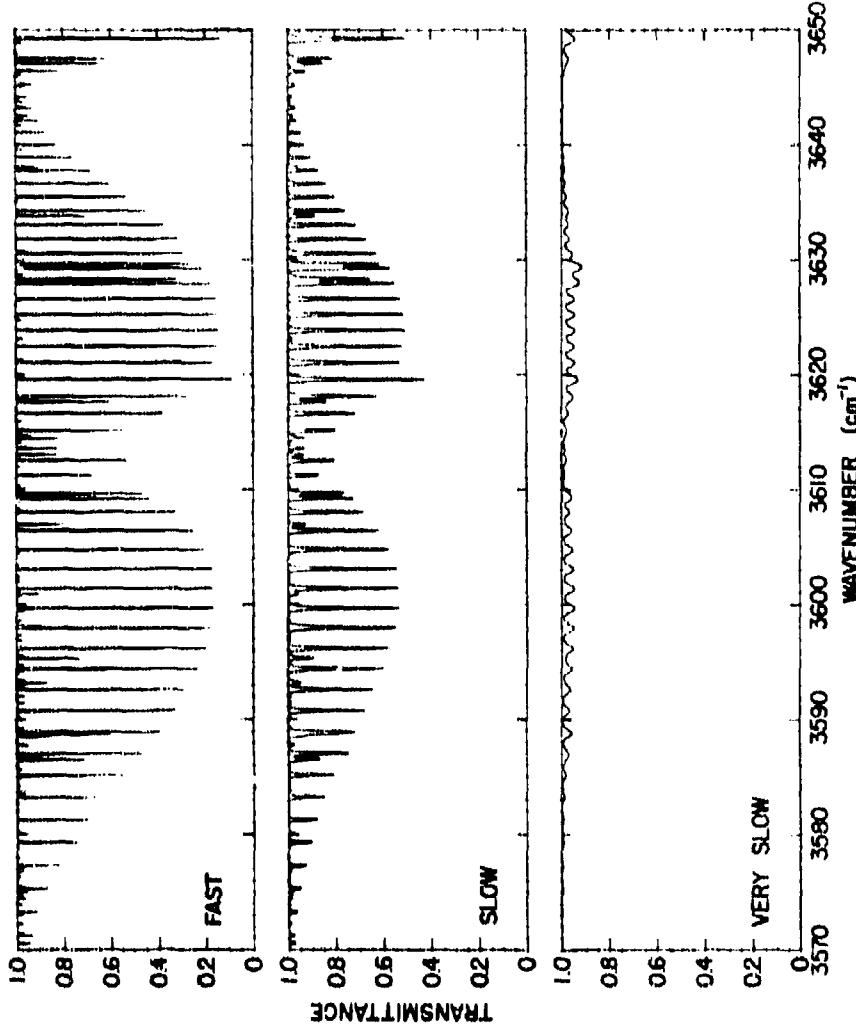


Figure 4. Results of the Convolution of the Three Individual Functions With the Spectral Line Data From 3570 cm^{-1} to 3650 cm^{-1} and a Sampling Interval $DV = 0.01\text{ cm}^{-1}$. From top to bottom the plots are the result of the fast function, slow function and very slow function convolutions respectively. The conditions are $P = 500\text{ millibars}$, $T = 240\text{ K}$, H_2O column density = 2.9×10^{18} , and CO_2 column density = $7.6 \times 10^{18}\text{ mol/cm}^2$. The results are expressed in transmission to increase the dynamic range of the presentation. The product of the three spectra give the total result shown in the upper spectrum of Figure 5.

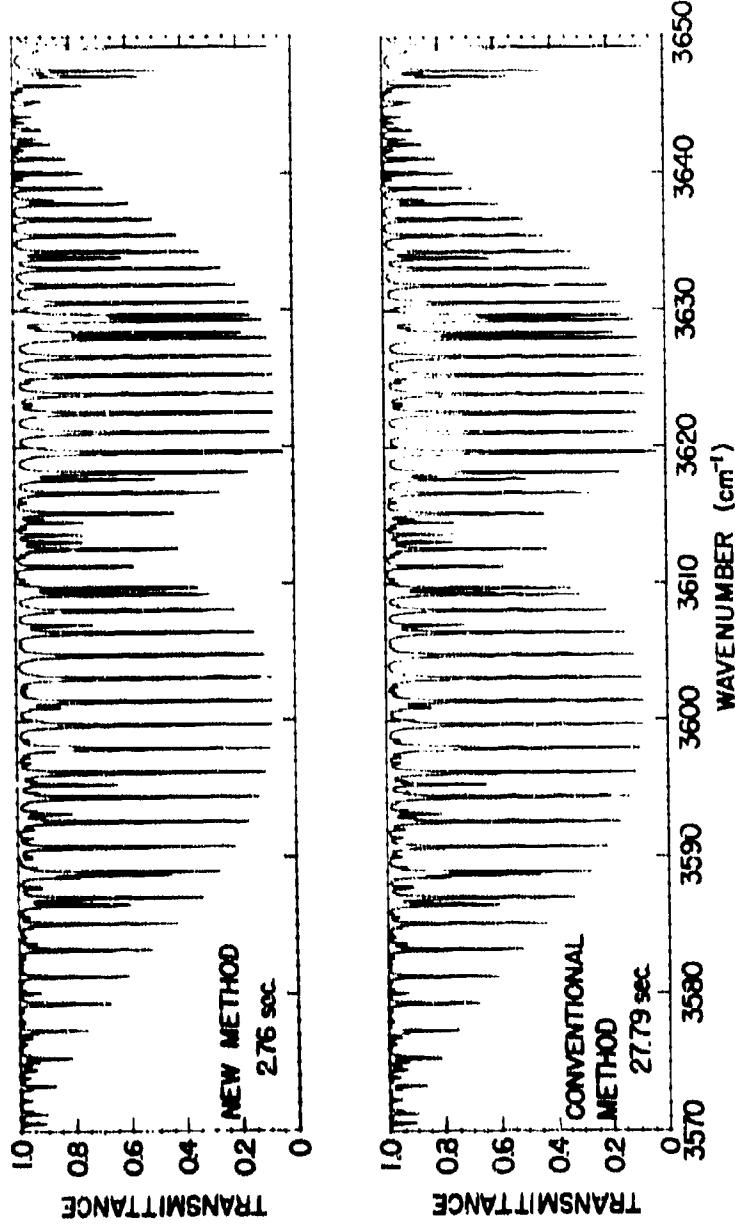


Figure 5. Spectral Transmission Results From 3570 cm^{-1} to 3650 cm^{-1} at a Sampling Interval of $\Delta V = 0.01\text{ cm}^{-1}$ for Conditions as in Figure 4: $P = 500$ millibars, $T = 240\text{ K}$, H_2O Column Density = 2.9×10^{18} , and CO_2 Column Density = $7.6 \times 10^{18}\text{ mol/cm}^2$. The upper trace is the result of the new algorithm and required 2.76 sec of computer time; the lower trace utilizing the conventional procedure required 27.79 sec. The times are for a CDC 6600 computer; read time has not been included. The AFGI line file was used for the spectral data from which 4401 lines were required.

In this expression, N_{tot} is the total number of halfwidths over which the function is to be computed, N_m is the number of halfwidths spanned by the m^{th} decomposed function, and $\delta x_1/\delta x_m$ is the ratio of the sampling interval of the first function to that required for the m^{th} function. In the present case for $N_{\text{tot}} = 48$, we have:

$$G = \frac{48}{(3 + 1 + 12 \cdot \frac{1}{4} + 48 \cdot \frac{1}{16})} = \frac{48}{9} = 5 \frac{1}{3} . \quad (42)$$

There are some constraints that are not explicit in the above expression; in particular, there are limitations on how small N_m can be and still have the function be reconstructable at the sampling interval. The algorithm provides greater improvement as the extent of the convolution is increased. For instance, if it were desired to reproduce the function over 192 halfwidths using four decomposed functions, a given choice would give a gain of

$$G = \frac{192}{(3 + 1 + 12 \cdot \frac{1}{4} + 48 \cdot \frac{1}{16} + 192 \cdot \frac{1}{64})} = \frac{192}{12} = 16 . \quad (43)$$

It must be emphasized that this choice, though workable, is not unique and other choices might provide greater improvement.

The algorithm, as it has been developed, must include protection against lines having halfwidths deviating significantly from the average. For anomalously wide lines, the convolution process can overflow the output arrays at the beginning and ending of the panels. In this case the line width is reset to the maximum width that can be treated: twice the average value. A series of + signs and the line identification are outputted when such a situation is encountered. In general, wide lines are not a problem since the widest lines do not vary significantly from the mean. This case also provides protection when a sampling interval, D_s , is chosen which is not consistent with the mean halfwidth of the lines being considered or when the line width is incorrect on the input data file.

There are a number of anomalously narrow lines that are encountered such as the high J , low K_a water lines of type measured by Eng et al.¹⁷ For such lines, that is, $a_i \ll \bar{a}$, where a_i is the width of the i^{th} line and \bar{a} is the mean width, not only does the algorithm fail, but depending on the phasing of the sampling interval, the contribution of the transition can be missed entirely. Such lines have been

17. Eng, R.S., Kelley, P.L., Mooradian, A., Calawa, A.R., and Harmon, T.C. (1973) Tunable laser measurements of water vapor transitions in the vicinity of $5 \mu\text{m}$, Chem. Phys. Letters 18:524.

treated by setting the halfwidth to the sampling interval, that is, $\alpha_1 = DV$. In this case, a series of - signs and the line identification are printed out. Although the convolution will be undersampled for such lines, information pertaining to the line is retained. The only strictly correct way to handle these narrow lines, is to choose a DV equal to $\alpha_1/4$ where α_1 is characteristic of the narrowest line encountered. Such a choice of DV may necessitate the use of the conventional convolution method, since the sampling interval is so small. Computational requirements may dictate the calculation of the absorption coefficient over a small wavenumber region around the lines in question. In general, these lines are relatively weak, and except in cases such as laser transmission problems the method used here proves to be satisfactory.

5. EXTENSIONS OF THE METHOD

As has been indicated in the previous section, the calculation of the far line wings may be readily incorporated into the program. However, for the reasons stated it is recommended that this far wing contribution be handled as a continuum contribution in the wavenumber domain. Similarly, the variation of line shape with molecular species may be incorporated into the algorithm. It is well known that the wings of air broadened CO₂ lines are "subLorentzian" and a very slow function specific to the CO₂ molecular species could be utilized in place of the Lorentz very slow function. However, there is no compelling evidence that the subLorentzian behavior occurs within 5 cm⁻¹ of the line center at atmospheric pressure so again, it is considered more appropriate to include the absorption due to the far wings of the CO₂ lines as a properly chosen continuum function. The best available theories indicate that the subLorentzian behavior is wavenumber dependent. Incorporation of this effect is not amenable to a direct extension of the algorithm since the argument of the Lorentz function is in terms of halfwidths. The wing modification may be readily added to the calculation by including the contribution into the very slow array before the three arrays are combined into the final output array. Continuum contributions of this type are important for CO₂ in spectral regions such as the 01001 band center where the wing contribution from the Q branch is not negligible as at 667.2 cm⁻¹ and beyond the band head from 780 cm⁻¹ to 890 cm⁻¹.

The convolution method described here offers even more significant computational savings for the multi-layer case as it is generally applied to atmospheric modeling. The sampling interval is determined for each layer based on the average halfwidth which is a function of temperature and pressure for the layer. The remaining problem is to merge the results of each layer using appropriate

interpolation techniques to obtain the transmission or emission for the case being considered. In the conventional approach to the problem, the calculations are performed at the finest sampling interval required and is the same for all layers; that is, the sampling interval required to perform the absorption coefficient calculation at 60 km is the same as that used at 0 km. This approach requires the calculation of many more output points than is necessary at altitudes below the original altitude with a proportional increase in computational effort. With the scheme proposed here, the computational time is the same for each layer.

6. LINE SHAPES OTHER THAN LORENTZIAN

The present method is clearly not restricted to any particular function such as the Lorentz function. The Doppler line shape may readily be tabulated either directly or as decomposed functions depending on the definition of halfwidth. For many atmospheric modeling cases, the line shape of greatest interest is the Voigt¹⁸ shape. Current efforts are being directed toward incorporating the Voigt line shape into the new algorithm. Preliminary indications are that the Voigt profile, computed to an accuracy similar to that attained for the Lorentz, will require only slightly more computer time than that required for the Lorentz function. Details of the procedure for calculating the Voigt shape will appear in a subsequent publication. Although the program assumes a symmetric line shape and only half the scanning function is stored, an asymmetric shape can be utilized by storing the entire function and making minor changes to the program.

7. THE PROGRAM

The program has been written to use line parameter input data consistent with that contained on the AFGL line parameter tape.¹ The line data is reformatted onto a binary file which contains the line data pertinent to the molecules and wave-number range of interest. This step has been taken to keep read time consistent with the time required to perform the calculational part of the program. The control parameters are read from the input file and written to the output file; "tape 3" is the binary file containing the line parameter data; and KKFILE is the binary output file. KKFILE contains a header record which includes the identification information, SECANT, temperature, pressure, molecular identification, and molecular column densities of the homogeneous layer. The first record for each output panel is a header record for the panel which contains the wavenumber values

18. Voigt, S. Münch, Ber. (1912) p. 803.

of the first and last absorption coefficient values of the panel, the wavenumber increment between output points and the number of output points. The second record of the panel contains the array of absorption coefficient values resulting from the convolution calculation. The current version of the program outputs a maximum of 2400 values; in general, the first and last panel are shorter.

The program consists of the main program HIRACC; subroutines SHAPE, MOLEC, RDFILE, CONVFN, and PANEL; and the function QVRFAC. The overall strategy of the program is indicated in Figure 6. All the subroutines are called from the main program and the flow of the program is easily traced. Subroutine SHAPE sets up the three convolution functions XF, XS, XVS used to define the Lorentz function from 0 to 48 halfwidths.

Subroutine MOLEC in conjunction with function QVRFAC, makes the molecular identifications with the line parameter file, and determines the correction factors for the line intensities (SCOR) and the halfwidths (ALFCOR). The quantity, SCOR, is the correction factor due to the temperature dependence of the vibrational and rotational partition sums. The vibrational partition sum is calculated for a given molecular type as

$$Q_v(T) = \prod_{i=1}^N \frac{1}{\left(1 - e^{-h\nu_i/kT}\right)d_i} \quad (44)$$

where ν_i is a fundamental vibrational frequency and d_i is the degeneracy of the vibration. The temperature dependence of the rotational partition sum is given by

$$\frac{Q_R(T_0)}{Q_R(T)} \approx \left(\frac{T_0}{T}\right)^F \quad (45)$$

where $F = 1$ for linear molecules and 1.5 for nonlinear molecules. The reference temperature, T_0 , is taken as 296 K, consistent with the AFGL Line Listing. For further discussion of these topics, see Herzberg¹⁹ pp 503 ff. The partition sum calculations are performed by QVRFAC and the necessary molecular parameters are contained in data statements in subroutine MOLEC. The quantity, ALFCOR, is the correction factor due to the pressure and temperature dependence of the collision broadened halfwidth. The temperature dependence of the halfwidth has

19. Herzberg, G. (1945) Molecular Spectra and Molecular Structure II. Infrared and Raman Spectra. D. Van Nostrand and Co., Princeton.

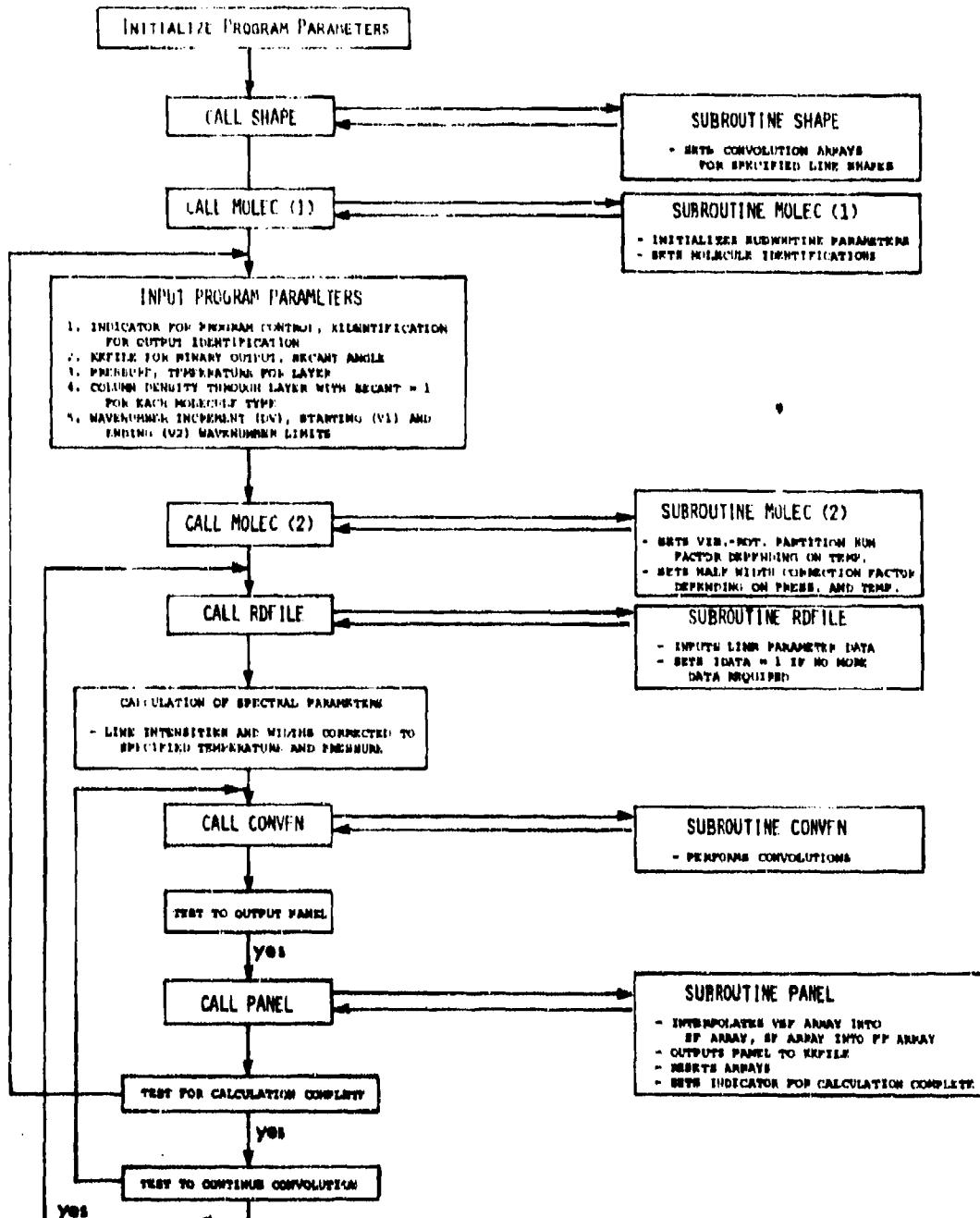


Figure 6. Flow Diagram for HIRACC Program

been taken as $(T_o/T)^{0.5}$ although calculations based on the Anderson - Tsao - Curnutte theory are reported to give a temperature dependence $\approx (T_o/T)^{0.75}$ (Varanasi²⁰).

Subroutine RDFILE reads the blocked binary line parameters over the wavenumber range for which line data is required. The line parameters include the wavenumber value of the transition (GNU, cm^{-1}), the intensity of the transition at 296^0K (S, $\text{cm}^{-1}/(\text{mol/cm}^2)$), the collision broadened halfwidth at half maximum for 296^0K and atm pressure (ALFAO, cm^{-1}), the lower state energy of the transition (EPP, cm^{-1}), and the molecule identification number (MOL). If the line parameter data is insufficient to complete the specified calculation, the message "end of file on disk" is printed on the output file. If no further line data is required IDATA is set to 1, and control is returned to the main program.

At this stage of the main program, an effective optical depth is calculated for each line which is dependent on the column density of the layer, the secant of the angle through the layer, the temperature of the layer, and the halfwidth of the line, $\sigma(\text{ALFI})$.

The effective depth, u' ,

$$u' = \left(\frac{1}{\alpha} \right) \cdot w \cdot \sec \cdot \left(\frac{Q_{VR}(T_o)}{Q_{VR}(T)} \right) \cdot \left[\exp \left(\frac{E''}{kT_o} - \frac{E''}{kT} \right) \right] \cdot \left[\frac{1 - \exp \left(- \frac{\hbar\nu}{kT} \right)}{1 - \exp \left(- \frac{\hbar\nu}{kT_o} \right)} \right] \quad (46)$$

where w is the absorber column density, E'' is the lower state energy, $Q_{VR} = Q_V Q_R$, and the other quantities have been previously defined. In terms of the program coding the effective depth appears as:

```
EFDEPTH=RECALF*W*SEC*SCOR*EXP(EPP/XKTFAC)*(1.-EXP(-GNU/TEMPO))/  
(1.-EXP(-GNU/TAVE)).
```

As previously discussed, the proper sampling interval, DV, should be 0.25 times the average line halfwidth. If the halfwidth, ALFI, is less than the sampling interval, the halfwidth is set to the sampling interval and a series of minus signs is written to the output file. If the halfwidth exceeds a maximum value (ALFMAX) where ALFMAX = BOUND/48 and BOUND is the maximum value in wavenumbers over which a line can be calculated, the halfwidth is reset to ALFMAX, and a series of + signs is written to the output file. Included in the records indicating the

20. Varanasi, P., and Ko, F.K. (1977) Intensity and transmission measurements in the ν_3 fundamental of N_2O at low temperatures, Thirty Second Symposium on Molecular Spectroscopy, Paper RF'5, Columbus, Ohio

resetting of the halfwidth is the wavenumber value of the transition (GNU), the intensity (S) and the halfwidth (ALFAO) values of the transition from the line parameter tape, the calculated value of the halfwidth (ALFI), the value to which the halfwidth (DV or ALFMAX), and molecular identification number (M). If the number of halfwidth changes (NCHNG) exceeds 100, the computation is terminated.

Subroutine CONVFN is a tightly written subprogram in which considerable effort has been taken to minimize operations in the DO 30 loop. This subroutine performs the triple convolution of X^F, XS, and XVS with the line data putting the results in the proper elements of FF, SF, and VSF respectively. A simplified flow diagram appears in Figure 7. Control indicator IPANEL is set to IDATA if the DO loop over the lines (40) is satisfied indicating whether a panel is complete or more lines are required. If the line DO loop (40) is not completed, IPANEL is set to 1 indicating that a complete panel has been calculated. Control is returned to the main program.

If IPANEL has been set to 1, subroutine PANEL is called. Subroutine PANEL performs a four point Lagrange interpolation of the VSF array into the SF array and the SF array into the FF array, thus combining the results of three independent convolutions into a final result. A general flow diagram of panel is given in Figure 8. Care is taken to store array values required for the interpolation of subsequent panels. VFT is the wavenumber value of the first element of the FF array, which is common to the first element of the SF and VSF arrays. A binary header record is written to the binary file (KKFILE) for each panel which includes the wavenumber value (VIP) of the first element of the panel (FF(NLO)), the wavenumber value (V2P) of the last element of the panel (FF(NHI)), the wavenumber increment (DV), and the number of absorption coefficient values outputted (NLIM). The second binary record contains the NLIM values of the absorption coefficient from the FF array. The arrays are appropriately shifted and reset in preparation for the computation of subsequent panels. Control is again returned to the main program, HIRACC.

After a panel is written to KKFILE, a record is written to the output file indicating the current value of the time, the time spent in RFILE, in CONVFN, and in PANEL (the units are seconds). Also included in this record are the first and last wavenumber values of the panel, the wavenumber increment, and the number of values in the panel. A second record is written to the output file indicating the average value of the halfwidth, the number of lines read since the last panel was completed, and the total number of lines read since the initiation of the convolution calculation. Control is returned to statement 10 if the calculation is complete, or to statement 140 to continue the convolution process.

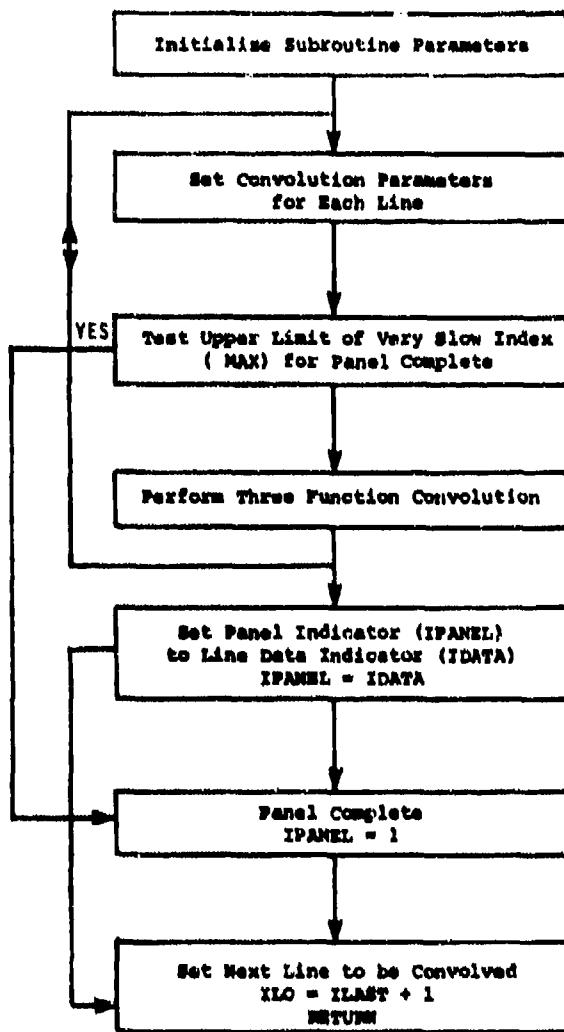


Figure 7. Flow Diagram for SUBROUTINE CONVFN

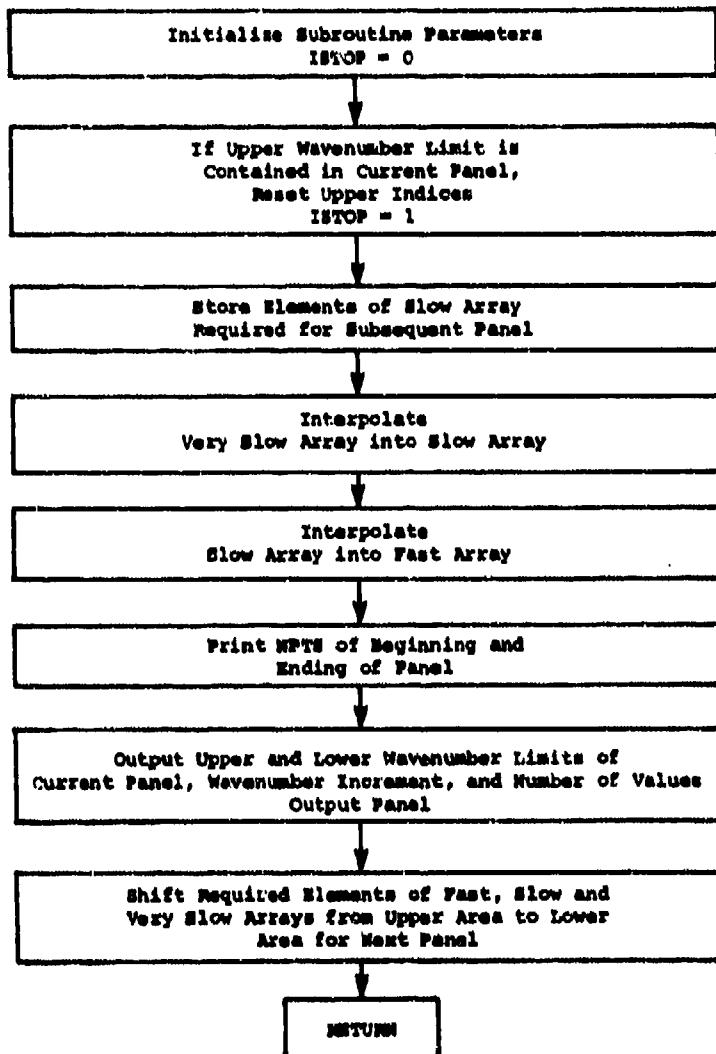


Figure 5. Flow Diagram for SUBROUTINE PANEL

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Appendix A

A listing of the Fortran program HIRACC is given in Table A1, together with the loader map, the input data, and the program output for the results which appear in Figure 5. In the output, it should be noted that one line is encountered that has a halfwidth which is larger than that allowed by the program. This line at $3558.4830 \text{ cm}^{-1}$ is an HDO line, molecular type 1, which has an erroneous halfwidth on the AFGL line tape used. The width for this line is reset to $\text{ALFMAX} = 0.08 \text{ cm}^{-1}$. Two H_2O lines occur at $3579.0810 \text{ cm}^{-1}$ which have anomalously narrow halfwidths. The values of these halfwidths are correct and are the high J, low K_A water lines of the type measured by Eng et al.¹ The widths for these lines are reset to $\text{DV} = 0.01 \text{ cm}^{-1}$. The beginning and ending sections of the output panels are printed out including the location in the FF array, the dimensionless absorption coefficient, and the wavenumber value associated with the respective elements of the FF array.

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Table A1. Listing of HIRACC Program

74/74 OPT=1

FTN 4.5+414

07/11/77

```

PROGRAM HIRACC(INPUT,CUTPUT,TAPE3,TAPE10)          000100
*                                             000110
*                                             000120
*                                             000130
*                                             000140
*                                             000150
*                                             000160
*                                             000170
*                                             000180
*                                             000190
*                                             000200
*                                             000210
*                                             000220
*                                             000230
*                                             000240
*                                             000250
*                                             000260
*                                             000270
*                                             000280
*                                             000290
*                                             000300
*                                             000310
*                                             000320
*                                             000330
*                                             000340
*                                             000350
*                                             000360
*                                             000370
*                                             000380
*                                             000390
*                                             000400
*                                             000410
*                                             000420
*                                             000430
*                                             000440
*                                             000450
*                                             000460
*                                             000470
*                                             000480
*                                             000490
*                                             000500
COMMON KK,KKFILE,PAVE,TAVE,U(?)          000510
COMMON GMU(250),S(250),ALFAG(250),ZPP(250),MOL(250)
COMMON FF(3200),SF(800),VSP(200)
COMMON /XSUB/ NCOP1,L1MIN,ILO,IHI,VTOP,V1,VE,DV,VFT,IOP1,IPAN
IEL,IData,ISTOP
COMMON /MXX/ NMF,NKF,NF,NMS,DXS,NS,NMV3,NFHAX,NFMAX,NVSH
1AX
COMMON /SUB1/ MAXF,MAXB,MAXVS,NLIMF,NLIMB,NLIVHS,NLO,NHI,DVS,DVS
COMMON /XPANEL/ V1P,VPP,NLIM,NLIMFT,NPTE
COMMON /XTIME/ TIME,TIMDF,TIMONV,TIMPNL
DIMENSION XID(1),NMOLID(7),N(7),SDR(7),ALFOOR(7)
DIMENSION XF(201),NS(201),XV(201)
DIMENSION EPDPTH(1), RECALF(1)
EQUIVALENCE (S,EPDPTH), (ALFA6,REGALF)
PIE,PIEIN(1,1)
RECPIN(1,1)
NPTSP=1
NPTSP=10
NMF=3
DXF=0.02
NPF=151
NMS=17
DXS=0.00
NMS=191
NMV3=48
DXVS=32
NMV3=191
NFMAX=201
NMV3=201
NMV3MAX=201
NMV3TTS=32
NLIMF=NLIMF/40+1
NLIMB=(NLIMB/40)+1
NOTE (DXVS/DXF) IS 16 AND (DXVS/DX) IS 4          000520
000530
000540
000550
000560
000570
000580
000590
000600
000610
000620
000630
000640
000650
000660
000670
000680
000690
000700
000710
000720
000730
000740
000750
000760
000770
000780
000790
000800
000810
000820
000830

```

Table A1. Listing of HIRACC Program (Cont.)

PROGRAM HIRACC			
75	NROUND=4R*(DXVS/DXF)	000840	
	MAXF=M*MF*BOUND	000850	
	MAYS=(MAXF/4)+1	000860	
	MAXVS=(MAXV/4)+1	000870	
	CALL SHARC(XF,YS,XVS)	000880	
80	CALL MHOLC(1,MHOLID,7,TEMPO,TAVE,P0,PAVE,SCOR,ALFCOR)	000890	
	LIMIN=290	000900	
10	PRINT 900	000910	
	TIMHDF=TIMCNV=TIHPNL=0.	000920	
	READ 902, IND,(XIO(I)),I=1,7	000930	
85	IF (IND(NFL)) GO TO 160	000940	
	PRINT 904, (XIO(I)),I=1,7	000950	
	REWIND 3	000960	
	TEOP00	000970	
	READ 905, KKFILE,SECANT	000980	
90	PRINT 910, KKFILE	000990	
	PRINT 910, SECANT	001000	
	READ 910, PAVE,TAVP	001010	
	PRINT 920, PAVE,TAVE	001020	
	READ 920, IN(M),M=1,7	001030	
95	PRINT 930, (MHOLID(M),M=1,7)	001040	
	READ 930, DV,V1,V2	001050	
	PRINT 940, DV,V1,V2	001060	
	IF (DV.LE.0.0) GO TO 160	001070	
	DVSM=(DXVS/DXF)*DV	001080	
100	DVVS=(DXVS/DXF)*DV	001090	
	ROUND=FLOAT(BOUND)*DV/2.	001100	
	PRINT 942, BOUND	001110	
	ALFMX=BOUND/4,	001120	
	NL0=MHOLID01	001130	
105	NHI=NL1MF+NSHIFT-1	001140	
	DO 90 I=1,MAXF	001150	
	FF(I)=0.	001160	
	DO 90 J=1,MAXF	001170	
90	SF(I)=0.	001180	
110	DO 70 I=1,MAXVS	001190	
70	VSF(I)=0.	001200	
	PRINT 900	001210	
C		001220	
	WRITE (KKFILE) (XIO(I),I=1,7),SECANT,PAVE,TAVE,(MHOLID(M),M=1,7),	001230	
115	1 (IN(M),M=1,7), DV,V1,V2	001240	
C	VFT=V1-2.*BOUND	001250	
	VROT=V1-BOUND	001260	
	VTOPAVE+ROUND	001270	
120	P0=1813.0	001280	
	TEMPB=96.0	001290	
	XKT0=0.6991*TEMPO	001300	
	XKT=.6.6991*TAVE	001310	
125	XKTFAC=(1./XKT0)-(1./XKT)	001320	
	CALL MHOLC(2,MHOLID,7,TEMPO,TAVE,P0,PAVE,SCOR,ALFCOR)	001330	
	DO 68 M=1,7	001340	
	U(M)*H(M)*SCOR(M)*SECANT	001350	
	LINCNT=0	001360	
130	NCHND=0	001370	
C	CONTINUE	001380	
	ICNT=0	001390	
	SUMHAL=0.	001400	
135	C	001410	
140	CONTINUE	001420	
C	CALL SECOND (TIME)	001430	
	IF ((IEOF,NE,0) GO TO 160	001440	
145	CALL MHFILE (NU,0,ALFA0,APP,MOL)	001450	
	CALL SECOND (TIME)	001460	
	THRDY=THRDY+(TIME-TIME)	001470	
C	IF ((IEOF,NE,0) GO TO 160	001480	
150	C	MODIFY LINE DATA FOR TEMPERATURE,PRESSURE, AND COLUMN DENSITY	001490
	DO 130 X=1,LO,1	001500	
	MHOLID(X)	001510	
	EFDT(X)=EC(X)*U(M)	001520	
155		001530	

Table A1. Listing of HIRACC Program (Cont.)

```

PROGRAM HIRACC
      IF (EFDPOTH(I).LE.0.) GO TO 130
      ICNT=ICNT+1
      ALFI=ALFA0(I)*ALFDOR(M)
      IF (ALFI.GE.0.0) GO TO 110
      PRINT 945, GNU(I),S(I),ALFA0(I),ALFI,DV,M
      ALFI=DV
      NCHNG=NCHNG+1
      IF (ALFI.LE.ALFMAX) GO TO 120
      PRINT 946, GNU(I),S(I),ALFA0(I),ALFI,ALFMAX,M
      ALFI=ALFMAX
      NOHNG=NOHNG+1
      120 CONTINUE
      SUMALF=SUMALF+ALFI
      RECALF(I)=I./ALFI
      EFDPOTH(I)=EFDPOTH(I)*EXP((PP(I)*XKTFAC)*RECALF(I)*
      & (1.-EXP(-GNU(I)/TAVE))/(1.-EXP(-GNU(I)/TEMPO)))
      130 CONTINUE
      IF (NCHNG.GT.100) GO TO 160
      140 CONTINUE
      170 C
      CALL CONVFN (GNU,EFDPOTH,RECALF,MOL,FF,SF,VSF,XF,XS,XVS)
      C
      IF (IPANEL.EQ.0) GO TO 180
      C
      190 CONTINUE
      C
      CALL PANEL (FF,SF,VSF,KWFILE)
      C
      CALL SECOND (TIME)
      PRINT 955, TIME,TIMRDF,TIMCHV,TIMPHL,V1P,V2P,DV,NCLM
      AVALF=SUMALF/FLDAT(ICNT)
      LINCNT=LINCNT+ICNT
      PRINT 960, AVALF,ICNT,LINCNT
      SUMALF=0.
      ICNT=0
      IF (ISTOP.NE.1) GO TO 140
      END FILE KWFILE
      GO TO 18
      C
      180 180 CONTINUE
      STOP
      C
      C
      C
      190 900 FORMAT (1H1)
      902 FORMAT (1X,F4.0)
      904 FORMAT (1X,F4.0)
      905 FORMAT (1X,F15.6)
      910 FORMAT ("D OUTPUT FILE NO. = " I5)
      200 915 FORMAT (" SECON = " F15.5)
      920 FORMAT ("0 (1PE10.3)")
      925 FORMAT ("0 PRECS(MA) = " F12.6/ "0 TEMP(K) = " F11.2)
      930 FORMAT (3H0 COLUMN DENSITY(MOLECULES/CM**2) "
      1 / /(5X,A6," = " 1P10.3) )
      205 935 FORMAT (F10.4)
      940 FORMAT ("0 DV(0H-1) = " F12.4/0 V1(0H-1) = " F12.4/0 V2(0H-1) = "
      1 F12.4)
      942 FORMAT (17H0 2*OUND(0H-1) = ,F8.4)
      945 FORMAT ("-----" F10.4,E14.3,3F10.6,I5)
      210 950 FORMAT ("-----" F10.4,E14.3,3F10.6,I5)
      955 FORMAT (1HB,10X,"TIME",11X,"READ",6X,"CONVOLUTION",10X,"PANEL",6X,
      10V1,13X,"DV",11X,"DV",6X,"NVALUES",/4F15.3,F15.4,F10.4,110)
      960 FORMAT ("0 AVERAGE WIDTH = " F10.5," NO. LINES = " I10,
      1 , " TOTAL NO. OF LINES=" I10//)
      215 END

```

Table A1. Listing of HIRACC Program (Cont.)

SUBROUTINE SHAPE	74/74 DM1=1	FTN 4.5+414	07/11/77
1	SUBROUTINE SHAPE (XF, NS, XVS)		002250
	COMMON /XSUN/ RECPI		JL2260
	COMMON /NXXX/ NHF, OXF, NF, NMS, OXS, NS, NHVS, DXV1, NXVS, NFMAX, NSMAX, NVSH		002270
	IAX		002280
	DIMENSION XF(1), NS(1), XVS(1)		JC2290
5	XLCRNZ(X2)=1./1.-X2		002300
	SFN(X2)=A1*B1*X1+C1*D1*X2*X2 ²		002310
	VSFN(X2)=A2*B2*X2+C2*D2*X2 ²		002320
	A0(20)=(1.+3.*20**4+3.*20**4)/(1.+20**2)**3		002330
	B0(20)=-(1.+3.*20**2)/(1.+20**2)**3		002340
	C0(20)=(1./1.+20**2)**3		002350
10	C HATCH AT Z1 HALFWIDTHS		002360
	Z1=A1		002370
	A1=A0(Z1)		002380
	B1=B0(Z1)		002390
	C1=C0(Z1)		002400
15	C HATCH AT Z2 HALFWIDTHS		002410
	Z2=A2		002420
	A2=A0(Z2)		002430
	B2=B0(Z2)		002440
	C2=C0(Z2)		002450
20	TOTAL=0.		002460
	DO 10 I=1,NFMAX		002470
	XF(I)=0.		002480
	XF(I)=RECPI*(XLCRNZ(0.)-SFN(0.))		002490
25	SUM=XF(1)		002500
	DO 20 JJ=2,NF		002510
	XFLOAT=(JJ-1)*DXF		002520
	X=XFX		002530
30	XF(JJ)=RECPI*(XLCRNZ(X2)-SFN(X2))		002540
	SUM=SUM+XF(JJ)*Z.		002550
	CONTINUE		002560
	XF=INF).		002570
	SUM=SUM*DXF		002580
35	TOTAL=TOTAL+SUM		002590
	DO 30 I=1,NSMAX		002600
	NS(I)=0.		002610
	NS(I)=PECPI*(SFN(0.1)-VSFN(0.1))		002620
40	SUM=NS(I)		002630
	NS1=FLOAT(NHF)/DXS+1.001		002640
	DO 40 JJ=2,NS1		002650
	X=FLOAT(JJ-1)*DXS		002660
	X=XFX		002670
	X(JJ)=RECPI*(SFN(X2)-VSFN(X2))		002680
45	SUM=SUM+X(JJ)*Z.		002690
	CONTINUE		002700
	NS1=NS1+1		002710
	DO 40 JJ=NS1,NS		002720
	XFLOAT=(JJ-1)*DXS		002730
50	X=XFX		002740
	X(JJ)=RECPI*(XLCRNZ(X2)-VSFN(X2))		002750
	SUM=SUM+X(JJ)*Z.		002760
	CONTINUE		002770
	NS1=NS1+0.		002780
	SUM=SUM*DXS		002790
55	TOTAL=TOTAL+SUM		002800
	DO 40 I=1,NVSMAX		002810
	NVS(I)=0.		002820
	NVS(I)=RECPI*VSFN(0.1)		002830
60	SUM=XVS(1)		002840
	NVS1=FLOAT(NHVS)/DXVS+1.001		002850
	DO 60 JJ=2,NVS1		002860
	XFLOAT=(JJ-1)*DXVS		002870
	X=XFX		002880
65	XV(JJ)=RECPI*VSFN(X2)		002890
	SUM=SUM+XV(JJ)*Z.		002900
70	CONTINUE		002910
	NVS1=NVS1+1		002920
	DO 60 JJ=NVS1,NVS		002930
	XFLOAT=(JJ-1)*DXVS		002940
	X=XFX		002950
75	XV(JJ)=RECPI*(XLCRNZ(X2))		002960
	SUM=SUM+XV(JJ)*Z.		002970
	CONTINUE		002980
	SUM=SUM*DXVS		002990
80	TOTAL=TOTAL+SUM		003000
	RETURN		003010
85	END		003020
			003030
			003040
			003050
			003060

Table A 1. Listing of HIRACC Program (Cont.)

SUBROUTINE MOLEC	74/74 OPT=1	FTN 4.4+414	07/11/77
1	SUBROUTINE MOLEC(IND,NMOLID,NMOLEO,TEMPO,TEMP,P0,P,SCON,ALFCOR)	003060	
2	DIMENSION NMOLID(1),SCON(1),ALFCOR(1)	003070	
3	DIMENSION H(7,4),ND(7,4)	003080	
4	COMMON /XMOLEO/ MM(7),MV(7),W1(7),W2(7),H2(7),	003090	
5	1, W3(7),M3(7),M4(7),DV(7),ROTFAC(7)	003100	
6	DATA MM(1),MV(1),W1(1),W1(1),W2(1),M3(1),M3(1),W4(1),W4(1),	003110	
7	1, DV(1),ROTFAC(1)/	003120	
8	1 6H 28 ,3, 3651.7,1, 1995.8,1, 3755.8,1, 0, ,0, 1.0004, 1.5 /	003130	
9	DATA MM(2),MV(2),W1(2),W2(2),W3(2),W4(2),M3(2),M4(2),	003140	
10	1, DV(2),ROTFAC(2)/	003150	
11	2 6H COE ,3, 1364.0,x, 687.3,2, 2349.3,1, 0, ,0, 1.0040, 1.0 /	003160	
12	DATA MM(3),MV(3),W1(3),W2(3),W3(3),W4(3),M3(3),M4(3),	003170	
13	1, DV(3),ROTFAC(3)/	003180	
14	3 6H O3 ,3, 1183.8,1, 780.9,1, 1042.1,1, 0, ,0, 1.0496, 1.5 /	003190	
15	DATA MM(4),MV(4),W1(4),W1(4),W2(4),W3(4),W4(4),M3(4),M4(4),	003200	
16	1, DV(4),ROTFAC(4)/	003210	
17	4 6H N2O ,3, 1284.9,1, 588.8,2, 2223.8,1, 0, ,0, 1.1267, 1.0 /	003220	
18	DATA MM(5),MV(5),W1(5),W1(5),W2(5),W3(5),W4(5),M3(5),M4(5),	003230	
19	1, DV(5),ROTFAC(5)/	003240	
20	5 6H CO ,1, 2143.3,1, 0, ,0, 0, ,0, 1.0280, 1.0 /	003250	
21	DATA MM(6),MV(6),W1(6),W1(6),W2(6),W3(6),W4(6),M3(6),M4(6),	003260	
22	1, DV(6),ROTFAC(6)/	003270	
23	6 6H CH4 ,4, 2914. ,1, 1533.3,2, 3018.9,3, 1297.9,3, 1.0866, 1.5 /	003280	
24	DATA MM(7),MV(7),W1(7),W1(7),W2(7),W3(7),W4(7),M3(7),M4(7),	003290	
25	1, DV(7),ROTFAC(7)/	003300	
26	7 6H O2 ,1, 1483.3,1, 0, ,0, 0, ,0, 1.0007, 1.0 /	003310	
27	C C MOLEC MAKES THE MOLECULAR IDENTIFICATIONS.	003320	
28	C C ALFCOR IS THE FACTOR BY WHICH THE COLLISION WIDTH MUST BE CHANGED	003330	
29	C C DUE TO THE DEPENDENCE ON PRESSURE AND TEMPERATURE	003340	
30	C C THE TEMPERATURE DEPENDENCE IS TAKEN AS (T/T0)^0.5	003350	
31	C C SCON IS THE FACTOR BY WHICH THE LINE INTENSITY IS CHANGED DUE TO	003360	
32	C C TEMPERATURE DEPENDENCE OF THE VIB AND ROT PARTITION SUMS	003370	
33	O O IF (IND,EQ.2) GO TO 20	003380	
34	NDIM=7	003390	
35	NVDIM=6	003400	
36	DO 10 M=1,NMOLEO	003410	
37	M(M,1)=W1(M)	003420	
38	M(M,2)=W2(M)	003430	
39	M(M,3)=W3(M)	003440	
40	M(M,4)=W4(M)	003450	
41	ND(M,1)=M1(M)	003460	
42	ND(M,2)=M2(M)	003470	
43	ND(M,3)=M3(M)	003480	
44	ND(M,4)=M4(M)	003490	
45	ND(M,5)=NM1(M)	003500	
46	ND(M,6)=NM2(M)	003510	
47	ND(M,7)=NM3(M)	003520	
48	10 NMOLID(M)=NM(M)	003530	
49	RETURN	003540	
50	CONTINUE	003550	
51	PRATIO=P/P0	003560	
52	TRATIO=TEMP/TEMP	003570	
53	DO 30 M=1,NMOLEO	003580	
54	SCON(M)=DV(1),TEMP,TRATIO,DV(1),ROTFAC,M,ND,NDIM,NVDIM	003590	
55	ALFCOR(M)=PRATIO^(TRATIO**0.5)	003600	
56	30 CONTINUE	003610	
57	RETURN	003620	
58	END	003630	

Table A 1. Listing of HIRACC Program (Cont.)

FUNCTION	74/74	OPT=1	FTN 4.5+614	07/11/77
QVRFAC				
1			FUNCTION QVRFAC(M,TEMP,TRATIO,QV0,ROTFAC,N,ND,MUIM,NVDIM)	003666
			DIMENSION QV0(1),ROTFAC(1)	003670
			DIMENSION N(MDIM,NVDIM),ND(MDIM,NVDIM)	003680
			QV=1.	003690
			DO 10 I=1,NVDIM	003700
			IF (ND(M,I),EQ,0) GO TO 20	003710
			SUM1=-EXP(-W(M,I)/TEMP)	003720
			IF (ND(M,I),GT,1) SUM1=SUM1*ND(M,I)	003730
	10		QV=QV/SUM1	003740
	20		QVRFAC=(QV0(M)/QV)*(TRATIO)**ROTFAC(M))	003750
			RETURN	003760
	0		END	003770
				003780
SUBROUTINE	74/74	OPT=1	FTN 4.5+614	07/11/77
RFILE				
1			SUBROUTINE RFILE (IINU,R,ALFAR,EPP,HOL)	003790
			COMMON /XBUD/ NEOF,I,LIMIN,ILO,IZH,VBOT,VTOP,VS,VB,UV,VFT,ZPAN	003800
			IEL,ZNAT,B,INTOP	003810
			DIMENSION IINU(1), B(1), ALFAR(1), EPP(1), HOL(1)	003820
			IDATA=0	003830
	10		READ (3) VMIN,VMAX,NREC	003840
			IF (EOF(3)) GO TO 20	003850
	20		IF (VMAX,GE,VBOT) GO TO 30	003860
			READ (3) VMIN	003870
	30		GO TO 10	003880
			READ (3) (IINU(I),B(I),ALFAR(I),EPP(I),HOL(I),I=1,NREC)	003890
			ILO=1	003900
			IMIN=MC	003910
			IF (IMIN,GE,VBOT) GO TO 50	003920
	40		DO 40 I=1,NREC	003930
			ILO=I	003940
			IF (IINU(I),GE,VBOT) GO TO 50	003950
	50		CONTINUE	003960
			50 IF (VMAX,LE,VTOP) RETURN	003970
	60		DO 60 I=1,NREC	003980
			IMIN=I	003990
			IF (IINU(I),GT,VTOP) GO TO 70	004000
	70		CONTINUE	004010
			70 IF (IMIN,LT,NREC) IDATA=1	004020
	80		RETURN	004030
			80 PRINT 900	004040
	90		IEOF=1	004050
			90 RETURN	004060
	0		004070	
	100		FORMAT 9 END OF FILE ON DISK*)	004080
			END	004090

Table A1. Listing of HIRACC Program (Cont.)

SUBROUTINE CONVENT	74/74 OPT+1	FTN 4.5+414	07/11/77	
1	SUBROUTINE DONVFN (GNU,FFDPTH,RECALF,MOL,PF,SF,VSF,XF,NS,XVS) COMMON /XSIMP/ RECPI,LIMIN,ILO,IMI,VBDT,VTOP,VI,V2,DV,VPT,TCDF,IPAN IEL,IData,ISTOP COMMON /WXXX/ NMF,OKF,NP,NMS,OKS,NS,NHVS,OKVS,NKVS,NFMX,NBMAX,NVSH 1AK COMMON /XUR1/ MAXF,MANS,MAXVS,NLIMF,NLIMS,NLIMVS,NL0,NHI,OVS,OVS COMMON /XTIME/ TIME,TIMHDF,TIMCHV,TIMPNL DIMENSION GNU(1),EVOPTH(1),RECALF(1),PF(1),SF(1) DIMENSION MOL(1) DIMENSION XF(1),XS(1),XVS(1),VSF(1)		004110 004120 004130 004140 004150 004160 004170 004180 004190 004200 004210 004220 004230 004240 004250 004260 004270 004280 004290 004300 004310 004320 004330 004340 004350 004360 004370 004380 004390 004400 004410 004420 004430 004440 004450 004460 004470 004480 004490 004500 004510 004520 004530 004540 004550 004560 004570 004580 004590 004600 004610 004620 004630 004640 004650 004660 004670 004680 004690 004700 004710 004720 004730 004740 004750 004760 004770 004780 004790 004800 004810 004820 004830 004840 004850 004860 004870 004880 004890 004900 004910 004920 004930 004940 004950 004960 004970 004980 004990 004999 END	004110 004120 004130 004140 004150 004160 004170 004180 004190 004200 004210 004220 004230 004240 004250 004260 004270 004280 004290 004300 004310 004320 004330 004340 004350 004360 004370 004380 004390 004400 004410 004420 004430 004440 004450 004460 004470 004480 004490 004500 004510 004520 004530 004540 004550 004560 004570 004580 004590 004600 004610 004620 004630 004640 004650 004660 004670 004680 004690 004700 004710 004720 004730 004740 004750 004760 004770 004780 004790 004800 004810 004820 004830 004840 004850 004860 004870 004880 004890 004900 004910 004920 004930 004940 004950 004960 004970 004980 004990 004999 END

Table A1. Listing of HIRACC Program (Cont.)

SUBROUTINE PANEL	74/74	OPTC1	FTN 4.5+4.6	07/11/77
1				004910
	SUBROUTINE PANEL (FFF, VF, VSF, KKFILE)			004920
	COMMON /XSU9/ RECPI, LIHIN, ILO, IHI, VBOT, VTOP, V1, V2, DV, VF1, IFOF, IPAN			004930
	IPL, IDATA, ISTOP			004940
	COMMON /SUB1/ MAXF, MAXS, MAXVS, NLINH, NLIMS, NLTHVS, NLO, NH1, DVVS, DVVS			004950
9	COMMON /XPAR1/ V1P, V2P, NLIN, NSHIFT, NPTS			004960
	COMMON /XTIME/ TIME, TIMDOP, TIMCHV, TIMPNL			004970
	DIMENSION FF(1), SF(1), VSF(1)			004980
	DIMENSION STOP(8)			004990
10	CALL SECOND (TIME)			005000
	X80=-7./120.			005010
	X61=100./120.			005020
	X82=35./120.			005030
	X100=-1./16.			005040
11	X11=9./16.			005050
	ISTOP=0			005060
	NHHI=(V2-VF1)/DV+1.5			005070
	IF (NHHI.GE.NHHI) ISTOP=1			005080
	IF (ISTOP.EQ.1) NH1=NHHI			005090
20	JMPR=NLO-NPTS			005100
	JJPR=NH1-NPTS			005110
	LIMLO=(NLO-1)/4-3			005120
	LIMHI=(NH1/4)+1			005130
21	C SLOW FUNCTION VALUES NEEDED FOR SUBSEQUENT PANELS ARE SAVED			005140
22	DO 10 J=1,8			005150
	SFSTOR(J)=SF(LIMHI+J-8)			005160
	DO 20 J=LIMLO,LIMHI,4			005170
	JVB=(J-1)/4+1			005180
	SF(J)=VF1(J)+VSF(JVB)			005190
30	SF(J+1)=SF(J+1)+X00*VSF(JVB-1)+X01*VSF(JVB)+X02*VSF(JVB+1)+X03*VSF			005200
	1(JVB+2)			005210
	+SF(J+2)*X10*(VSF(JVB-1)+VSF(JVB+2))+X11*(VSF(JVB)+VSF(JVB+1))			005220
	+SF(J+3)=SF(J+3)+X02*VSF(JVB-1)+X03*VSF(JVB)+X01*VSF(JVB+1)+X00*VSF			005230
35	1(JVB+2)			005240
	CONTINUE			005250
	DO 30 J=NLO,NHI,4			005260
	JB=(J-1)/4+1			005270
	FF(J)=VF1(J)+SF(JB)			005280
40	FF(J+1)=FF(J+1)+X00*SF(JB-1)+X01*SF(JB)+X02*SF(JB+1)+X03*SF(JB+2)			005290
	+SF(J+2)*X10*(SF(JB-1)+SF(JB+2))+X11*(SF(JB)+SF(JB+1))			005300
	+SF(J+3)=SF(J+3)+X02*SF(JB-1)+X03*SF(JB)+X01*SF(JB+1)+X00*SF(JB+2)			005310
45	CONTINUE			005320
	IF (NPTS.EQ.8) GO TO 60			005330
46	DO 40 J=NLO,JMPPR			005340
	VPOVF=FLOAT(J-1)*DV			005350
47	PRINT 900, J, FF(J), VF			005360
	DO 50 J=JMPR,NHI			005370
	VPOVF=FLOAT(J-1)*DV			005380
50	PRINT 900, J, FF(J), VF			005390
55	CONTINUE			005400
	DO 70 J=1,8			005410
	SF(LIMHI+J-8)=SFSTOR(J)			005420
	NLIM=NH1-NLO+1			005430
56	V1P=VF1+FLOAT(NLO-1)*DV			005440
	V2P=VF1+FLOAT(NHI-1)*DV			005450
	C VF1 IS FIRST FREQ OF PANEL			005460
60	C VF1 IS LAST FREQ OF PANEL			005470
	WRITE (KKFILE) V1P, V2P, DV, NLIN			005480
	WRITE (KKFILE) (FF(J),J=NLO,NHI)			005490
	VPOVF=FLOAT(NLIM-1)*DV			005500
65	IF (ISTOP.EQ.1) GO TO 140			005510
	JFA=1			005520
	DO 80 J=NHLIM,MAXF			005530
	FF(J)=VF1(J)			005540
66	JF=JFA			005550
	DO 90 J=JF,MAXF			005560
67	FF(J)=0.			005570
	JFA=1			005580
70	DO 100 J=NHLIMS,MAXS			005590
	SF(J)=VF1(J)			005600
71	JSA=JFA+1			005610
	DO 110 J=JSA,MAXS			005620
	SF(J)=0.			005630
72	JVB=1			005640
	DO 120 J=NHLIMVS,MAXVS			005650
	VSF(J)=VF1(J)			005660
73	JVB=JVB+1			005670
	DO 130 J=NHLIMVS,MAXVS			005680
	VSF(J)=0.			005690
74	NLO=NH1+1			005700
	CALL SECOND (TIME)			005710
	TIMPNL=TIMPNL+TIME-TIME1			005720
	RETURN			005730
75	O			005740
	C			005750
	C			005760
76	FORMAT (110,24X,112.5,F12.5)			005770
	END			005780

Table A2. Loader Map for HIRACC Program

LOAD MAP - 4194ACC		CYBER LOADER 1.1-024		07/11/77 21:37:23.		PAGE 1	
FILE OF THE LOAD	LEN OF THE LOAD	111	36				
TRANSFER ADDRESS --- HI-FACT	43546						
PROGRAM AND BLOCK ASSIGNMENTS.							
BLOCK	ADDRESS	LENGTH	FILE	DATE	PROCESS PER LEVEL	HARDWARE	COMMENTS
/5500/	111	36					
PER1/	127	14					
/5091/	163	12					
XPANEL/	155	5					
/XTIME/	162	6					
WIRFACT	166	12686	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
SHPPE	12572	252	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
SWOLEA/	13855	124	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
WIFEC	13283	228	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
STIFAC	13421	211	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
STIFIL	13532	164	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
CONFVN	13716	238	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
PANEL	14246	657	L60	37/11/77 FTN	0-5 414	666x 1	CPT=1
/FCL-C/	16635	23					
/FCL-10/	16660	131					
CART10-	15811	44	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	COMMON CODED I/O SUBROUTINES AND UTILITIES.
DATA10-	15711	6	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	FCL INITIALIZATION ROUTINE.
ENDFIL-	15975	61	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	WRITE END OF LOGICAL FILE MARK.
FECHKE	15156	41	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	INITIALIZE CONSTANTS.
FLYGRTS	15217	316	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	COMMON FLOATING OUTPUT CODE.
FOREST-	15527	621	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	FOOTBALL OBJECT LIBRARY UTILITIES.
INC-DIM	16418	277	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	COMMON INPUT FORMATING CODE.
TOPIC-	16627	162	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	FORMAT TO RAN FORTRAN K-CODE.
RODER-	17407	456	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	OUTPUT FORMAT INTERFACE.
GUTC1-	17465	172	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	FORMATTED WRITE PRIVATE RECORDS.
ALG	17657	72	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	COMMON AND NATURAL LOGARITHMIC EXPONENTIAL FUNCTIONS. E TO POWER X. PI = 3.141592653589793.
EXP	17752	75	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	LINK BETWEEN SYSTEM AND INITIALIZED FILES.
SYSAIDS-	21247	1	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	TEST FOR END OF FILE STATUS.
EOP	21358	16	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	COMMON FLOATING INPUT CONNECTIONS.
FLYIN-	23066	154	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	CRACK APILIST AND FORMAT FOR KCODE/K-CODE.
FWAPL-	26262	352	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	FCL PMSL UTILITIES.
FORUTIL-	28616	16	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	LOCATE AND FIT GIVEN A FILE NAME.
CFITF1-	28632	42	SL-FORTRAN	38/11/76 COMPASS	3+	2-414	

Table A2. Loader Map for HIRACC Program (Cont.)

LDR MAP - HIRACC			
/DB,OUT,F/	20574	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
TAPE,	21123	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
DATA,F/	486	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
SET,IN	27845	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
OUTCNS	22251	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
REALMS	22424	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
CLOCK	22463	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
ASIN	22514	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
SYS=151	22645	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
XDIS	22667	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
RTDIS	22677	SL-FORTTRAN	16/11/76 COMPASS 3. 2-41
ROTBL,IN	22736	SL-SYSTIC	15/28/76 COMPASS 3. 75125
CIB,IN	22714	SL-SYSTIC	15/28/76 COMPASS 3. 75125
/ABL,IN/	18	SL-SYSTIC	15/28/76 COMPASS 3. 75125
MOVE,IN	22754	SL-SYSTIC	15/28/76 COMPASS 3. 75125
NET,IN	23050	SL-SYSTIC	15/28/76 COMPASS 3. 75125
JMPB,IN	23153	SL-SYSTIC	15/28/76 COMPASS 3. 75125
OPEN,IN	23314	SL-SYSTIC	15/28/76 COMPASS 3. 75125
OPEN,F/	23117	SL-SYSTIC	15/28/76 COMPASS 3. 75125
OPEN,IN,F/	23323	SL-SYSTIC	15/28/76 COMPASS 3. 75125
OPEN,IN,F/	23327	SL-SYSTIC	15/28/76 COMPASS 3. 75125
STEM,IN,F/	23564	SL-SYSTIC	15/28/76 COMPASS 3. 75125
STEM,IN,F/	23565	SL-SYSTIC	15/28/76 COMPASS 3. 75125
PUT,SO	23574	SL-SYSTIC	15/28/76 COMPASS 3. 75125
WR,SP	25156	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SL,SP,FLV	25476	SL-SYSTIC	15/28/76 COMPASS 3. 75125
CLOC,IN	25465	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SET,IN,F/	25478	SL-SYSTIC	15/28/76 COMPASS 3. 75125
QTTR,SIQ	25479	SL-SYSTIC	15/28/76 COMPASS 3. 75125
WEND,SIQ	25511	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SCPL,FLV	25753	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SDF,L,SIQ	25762	SL-SYSTIC	15/28/76 COMPASS 3. 75125
ERR,IN	26631	SL-SYSTIC	15/28/76 COMPASS 3. 75125
CMR,B,SIQ	26445	SL-SYSTIC	15/28/76 COMPASS 3. 75125
BSJL,B,SIQ	26446	SL-SYSTIC	15/28/76 COMPASS 3. 75125
OPEN,SIQ	26531	SL-SYSTIC	15/28/76 COMPASS 3. 75125
ONE,L,SIQ	27112	SL-SYSTIC	15/28/76 COMPASS 3. 75125
/SET,IN/	27327	SL-SYSTIC	15/28/76 COMPASS 3. 75125
PLEB,IN	27864	SL-SYSTIC	15/28/76 COMPASS 3. 75125
ELSF,SIQ	27192	SL-SYSTIC	15/28/76 COMPASS 3. 75125
AC,SV,FLV	27234	SL-SYSTIC	15/28/76 COMPASS 3. 75125
QSL,SIQ	27242	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SEN,F/	27356	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SEN,SIQ	27375	SL-SYSTIC	15/28/76 COMPASS 3. 75125
SET,FO/F	27426		

Table A2. Loader Map for HIRACC Program (Cont.)

LOAD MAP - MESSAGES	
1001-017	27435
007-028	27445
258	3085
30543	101
30544	98
30545	106
30546	37
30547	3021
	12525

.721 CP SECONDS

641000 CR STORAGE USED

71 TABLE MODES

PROCESSES SYSTEM REQUEST,

Table A3. Sample Output From HIRACC Program

OCD AND HOW = 3550 TO 3650 FOR REPORT

OUTPUT FILE NO. = 18

SECANT = 1.00000

PRESS (PSI) = 500.00000

TEMP(1) = 248.00

COLUMN SENSITIVITY (MOLECULES/CHNO2)

H2O = 2.910E+18

CO2 = 7.688E+18

O3 = -8.

HCNO = -8.

CO = -8.

CH4 = -8.

O2 = -8.

DV(CH-1) = .0100

V1(CH-1) = 3550.0000

V2(CH-1) = 3650.0000

ZERO DNU(CH-1) = 3.5625

Table A3. Sample Output From HIRACC Program (Cont.)

*****	3558.6738	-1.69E-47	-1.5899E-02	.886618	.886606	1
769	-1.9788E-02	3552.03055				
770	-1.6811E-02	3553.03106				
771	-1.7869E-02	3553.03208				
772	-1.6573E-02	3553.03105				
773	-1.6692E-02	3553.03439				
774	-1.7611E-02	3553.05036				
775	-1.6789E-02	3553.06600				
776	-2.8963E-02	3553.08735				
777	-2.4889E-02	3553.08480				
778	-2.6353E-02	3553.03925				
779	-1.6628E-02	3553.01988				
780	-4.1268E-02	3553.01187				
781	-4.6754E-02	3553.01238				
782	-4.8183E-02	3553.01308				
783	-4.6547E-02	3553.01468				
784	-3.8362E-02	3553.01501				
785	-3.2832E-02	3553.01188				
2416	-5.6406E-01	3566.04708				
2617	-6.6591E-01	3566.06466				
2418	-7.8283E-01	3566.04988				
2419	-9.6324E-01	3566.05386				
2420	-1.1358E+00	3566.05308				
2421	-1.3222E+00	3566.05208				
2422	-1.5281E+00	3566.05388				
2423	-1.6691E+00	3566.05498				
2424	-1.7253E+00	3566.05508				
2425	-1.6592E+00	3566.05608				
2426	-1.5163E+00	3566.05708				
2427	-1.3229E+00	3566.05808				
2428	-1.1516E+00	3566.05908				
2429	-9.8761E-01	3566.06008				
2430	-8.7671E-01	3566.06108				
2431	-7.7873E-01	3566.06208				
2432	-7.0855E-01	3566.06308				
			CONVOLUTION	PANEL		
				-167		
TIME	READ					
5.754	-740					
AVERAGE WIDTH =	0.038668	NO. LINES =	1283	TOTAL NO. OF LINES =	1263	
				V ₂	3566.0300	
				V ₁	3551.0000	
				V _{1J,L}		
				VALUES		166*

Table A3. Sample Output From HIRACC Program (Cont.)

	TIME	READ	PANEL	V1	V2	DY	DY	N VALUES
33	3579.9818	-131E-97	-8121756	-9106441	-9106681	1	1	2444
34	3579.9871	-391E-97	3566.64988	3562038	356578	.3516686	1	2444
35		-562322E-91	3566.65839					
36		-52345E-91	3566.66988					
37		-62964E-91	3566.67388					
38		-66562E-91	3566.68888					
39		-73269E-91	3566.69388					
40		-84911E-91	3566.70393					
41		-10070E+00	3566.71888					
42		-12546E+00	3566.72388					
43		-16105E+00	3566.73888					
44		-19919E+00	3566.74388					
45		-22946E+00	3566.75882					
46		-22841E+00	3565.76088					
47		-19869E+00	3566.77888					
48		-15594E+00	3566.79888					
49		-11633E+00	3566.79980					
50		-91763E-91	3566.80988					
51		-28374E-91	3590.47980					
52		-29933E-91	3590.49988					
53		-32079E-91	3590.49990					
54		-34682E-91	3590.50988					
55		-37426E-91	3590.51888					
56		-48465E-91	3590.52088					
57		-63434E-91	3590.53888					
58		-46314E-01	3590.54888					
59		-49843E-01	3590.55888					
60		-53757E-01	3590.56888					
61		-58293E-01	3590.57988					
62		-64886E-01	3590.59388					
63		-74955E-01	3590.59988					
64		-79524E-01	3590.60888					
65		-90559E-01	3590.61888					
66		-10363E+00	3590.62988					
67		-11919E+00	3590.63788					
NO. LINES =	1000	TOTAL NO. OF LINES =	2263					
AVERAGE WIDTH =	6.638	READ	CONVOLUTION	1.253	1.253	1.253	1.253	2444

Table A.3. Sample Output From BURCC Program (Cont.)

Table A3. Sample Output From HIRACC Program (Cont.)

		AVERAGE WIDTH = .33852	NO. LINES = 1000	NO. OF LINES = 6633
32				
34				
35				
36				
37				
38				
39				
40				
41				
42				
43				
44				
45				
46				
47				
48				
49				
2416				
2417				
2418				
2419				
2420				
2421				
2422				
2423				
2424				
2425				
2426				
2427				
2428				
2429				
2430				
2431				
2432				
	TIME 0.24s	READ CONVOLUTION .827	PART • 313	V1 3614.640C
		2.222		V2 3636.637G
				DV .610C
				VALUES 2432

Table A.3. Sample Output From HIRACC Program (Cont.)

Table A4. Data Cards for HIRACC Program

1 020 AND NOH = 3558 TO 3656 FOR REPORT
18 1.
588.8E 86 248. E 86
2.988E 18 7.686E 16 0.888E 08 0.888E 05 0.888E 02 0.888E 00 0.888E 00
.91 3558. 3656.
99999
802

Appendix B

A listing of a Fortran program to plot the results from HIRACC in transmission is given in Appendix B. This program, entitled TPLOT, was written for the CDC 6600 and the plotting functions are specific to that computer. This plot program is consistent with the sampling criterion and the interpolation methods described in this report. The four point Lagrange interpolation is used to generate three intermediate points for each output interval of HIRACC. The listing of the program, the output, and the input data, are given in Table B1. Two data cards are required: the first contains 30 characters of identification, and the second contains the upper and lower limits (cm^{-1}) of the wavenumber range and the distance in inches (DX) over which this wavenumber range is to be plotted.

Table B1. Listing of TPLOT Program

PROGRAM TPLOT	74/74 OPT=1	FTN 4.5+414	07/12/77
1	PROGRAM TPLOT(INPUT,OUTPUT,TAPE1)		000100
	DIMENSION PROGID(3)		000110
	COMMON Y(2500),XP(2500),YP(2500)		000120
	DIMENSION XID(7),W(7),MHOL(7)		000130
	HEQD1H		000140
4	READ 910, PROGID		000150
	REWIND 10		000160
	PRINT 910, PROGID		000170
	CALL PLTIDS (PROGID,JUL,0,11,0,0,0)		000180
10	READ 970, VI,VZ,XSIZE		000190
	IF (XSIZE.LT.1,0E-5) GO TO 190		000200
	PRINT 930, VI,VZ,XSIZE		000210
	READ (10) (YTDX),I(1,7),SEQ,PO,TO,(MHOL(I),I=1,7),		000220
	(W(I),I=1,7),VUV,VIV,V2V		000230
15	DELY=0.5		J00240
	YT=10.		000250
	YT=YT-DELY		000260
	CALL SYMBOL(0.0,YT,.15,XID, 0.0,70)		000270
	YT=YT-DELY		000280
20	CALL SYMBOL(0.0,YT,0.15,91SEQ =,0.0,5)		000290
	CALL NUMBER(1.0,YT,0.15,SEQ,0.0,1)		000300
	YT=YT-DELY		000310
	CALL SYMBOL(0.0,YT,0.15,LHMPRESS (4R)=,0.0,11)		000320
	CALL NUMBER(2.0,YT,0.15,P9,0.0,2)		000330
25	YT=YT-DELY		000340
	CALL SYMBOL(0.0,YT,0.15,BHTEMP (K)= ,0.0,0)		000350
	CALL NUMBER(1.5,YT,0.15,T0, 0.0,1)		000360
	YT=YT-DELY		000370
	CALL SYMBOL(0.0,YT,0.15,20HAMOUNTS (MHOL/CH**2) =,0.0,20)		000380
30	DO 20 MH=1,7		000390
	YT=YT-DELY		J00400
	MH=0,		000410
	HIP=0,		J00420
	IF I(MH).LT.1,0E-33) GO TO 10		000430
35	IP=ALOG10(CH(MH))		000440
	HIP=IP		000450
	RHMH(MH)/10,++IP		000460
40	CALL SYMOL(0.0,YT,.15,HMOL(MH),0.0,0)		000470
	CALL SYMBOL(1.0,YT,0.15,HEQ,0.0,1)		000480
	CALL NUMBER(1.3,YT,0.15,4N,0.0,3)		000490
	CALL NUMBER(2.3,YT,0.15,HIP,0.0,-1)		000500
45	CONTINUE		000510
	YT=YT-DELY		000520
	CALL SYMBOL(0.0,YT,.15,10HUV(CH-1)= ,0.0,10)		000530
	CALL NUMBER(1.5,YT,0.15,VUV, 0.0,3)		000540
50	YT=YT-DELY		000550
	CALL SYMBOL(0.0,YT,0.15,10HV1(CH-1) = ,0.0,10)		000560
	CALL NUMBER(1.5,YT,0.15,VIV, 0.0,2)		000570
	YT=YT-DELY		000580
55	CALL SYMBOL(0.0,YT,0.15,10HV2(CH-1) = ,0.0,10)		000590
	CALL NUMBER(1.5,YT,0.15,VRV,0.0,2)		000600
	YT=7.0		000610
	CALL PLOT(1T,0.0,-J)		000620
60	DX=(VR-V1)/XSIZEx		000630
	CALL AXIS (0.0,0.0,1M ,1,10.0,40.0,0.0,0.1,20.0)		000640
	CALL AXIS (XBIZE,0.0,1M ,-1,10.0,40.0,0.0,0.1,20.0)		000650
	CALL AXIS (0.0,10.0,1M ,1,XSIZE,1.3,VI,EY,20.0)		000660
	CALL AXIS (0.0,0.0,10HAWENNUMBER,-10,XSIZE,0.0,VI,DX,20.0)		000670
65	READ (10) V1P,V2P,DV,NLTH		000680
	IF (EOF(11)) 100,60		000690
70	60 LIMHI=NLTH		000700
	IF (V1P.GT.V2P) GO TO 190		000710
	PRINT 920, V1P,V2P,DV,NLTH,LIMHI		000720
	READ (10) (V(I)),I=6,LIMHI		000730
	IC (V2P,LT,V1) GO TO 50		000740
75	ILO=3		000750
	IMH=ILU+600		000760
	DO 70 IC=4,LIMHI		000770
	GAV=V(I)		000780
	V(I)=0.		000790
80	IF (GAV,LT,20.0) V(I)=EXP(-GAV)		000800

Table B1. Listing of TPLOT Program (Cont.)

PROGRAM LINE		
	CONTINUE	000010
	IF (V1P.GT.V1) GO TO 80	000050
	VLO=(V1-V1P)/DV+5.001	000060
75	IMH=ILO+600	000070
	IF (V1P.LT.V2) GO TO 90	000080
	LIMMH=(V2-V1P)/DV+5.001	000090
	IF (LIMMH.LT.600) IMH=LIMMH-1	0000910
	CALL LPP (ILO,IMH,V1P,V1,DX,DV)	0000920
80	ILO=IMH	0000930
	IMH=ILO+600	0000940
	IF (IMH.LT.LIMMH) GO TO 90	0000950
	IMH=LIMMH-1	0000960
85	IF (ILO.LT.(LIMMH-1)) GO TO 90	0000970
	V(1)=Y(LIMMH-3)	0000980
	V(2)=Y(LIMMH-2)	001000
	V(3)=Y(LIMMH-1)	001010
	V(4)=Y(LIMMH)	001020
90	IF (LIMMH.LT.NLIM) GO TO 100	001030
	GO TO 90	001040
	CALL PLOT (XSIZE+3.0,0.0,-3)	001050
	CALL RNDPLT	001060
	STOP	001070
95	C	001080
	O	001090
960	FORMAT (E(1A2,F10.6))	001100
918	FORMAT (3A10)	001110
928	FORMAT (3F12.3,2E10)	001120
938	FORMAT (4F10.3,4E9)	001130
	END	

SUBROUTINE LPP

1	SUBROUTINE LPP (ILO,IMH,V1P,V1,DX,DV)	001140	
	COMMON V(2500),XP(2500),VP(2500)	001150	
	PR2NT 810, ILO,IMH,V1P	001160	
	DV1=DV/41	001170	
	DV2=DV1*DV1	001180	
	DV3=DV2*DV1	001190	
	X00=-7./120.	001200	
	X01=-65./120.	001210	
	X02=-5./120.	001220	
	X03=-4./120.	001230	
	X10=-1./16.	001240	
	X11=-9./16.	001250	
	IP=1	001260	
	IMAX=IMH-1	001270	
10	DO 10 I=ILO,IMAX	001280	
	XP(IP)=V1P+DV*FLOAT(I-5)	001290	
	XP(IP+1)=XP(IP)+DV1	001300	
	XP(IP+2)=XP(IP)+DV2	001310	
	XP(IP+3)=XP(IP)+DV3	001320	
	VP(IP)=V1	001330	
	VP(IP+1)=X00*Y(I-1)+X01*Y(I)+X02*Y(I+1)+X03*Y(I+2)	001340	
	VP(IP+2)=X10*Y(I-1)+X11*Y(I)+X01*Y(I+1)+X00*Y(I+2)	001350	
	VP(IP+3)=X00*Y(I-1)+X01*Y(I)+X02*Y(I+1)+X03*Y(I+2)	001360	
	IP=IP+6	001370	
15	10	CONTINUE	001380
	XP(IP)=V1P+DV*FLOAT(IMH-5)	001390	
	VP(IP)=Y(IMH)	001400	
	CALL LINE (XP,VP,IP,1,0,1,V1,DX,E,0.0+1,0.0)	001410	
	RETURN	001420	
20	O	001430	
908	FORMAT (E(1A2,F10.6))	001440	
918	FORMAT (2E10,F4E3,S,I,ILO,IMH,V1P)	001450	
	END		

Table B2. Loader Map for TPLOT Program

LINK MAP - TPLOT		FILE OF THE LOAD LBN OF THE LOAD		TRANSFER ADDRESS -- TPLOT		CROSS-REF FOR 1.1-428		17/12/77		JU.31.45.		PAGE 1	
TPLOT	111	TP177	160	177	160	07/12/77	FTN	4.5	414	665X	I	OPT=1	
LPA	7313	282	6	7312	6	07/12/77	FTN	4.5	414	665X	I	OPT=1	
PROGRAM AND BLOCK ASSIGNMENTS.													
BLK#	ADDRESS	LENGTH	FILE	DATE	PROCESSOR LEVEL	Hardware	Comments						
TPLOT	111	7177	160	07/12/77	FTN	4.5	414	665X	I	665X	I	OPT=2	
LPA	7313	282	6	7312	6	07/12/77	FTN	4.5	414	665X	I	OPT=2	
MAIN/	7521	1											
FC1-EDS/	7521	1											
STRICTY/	7522	1											
SA WORK/	7523	5	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
PLTIB3	7528	171	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
BUDRIT	7723	54	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
PLT01	7775	325	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
EXPLT	18322	235	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
FACT02	18597	16	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
SYMBOL	18575	387	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
PLT00	11360	432	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
PLCNET	11536	625	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
EMPLT	12163	168	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
12343	41	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2		
PLT02	12464	27	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
NAME	12627	52	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
CHELT	12501	11	UL-PEN	18/06/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
LATE	12512	318	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
NEFEEF	13822	233	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
NEUPEN	13255	66	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
PLOTS	13367	75	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
BUFF	13448	467	UL-PEN	18/06/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
ASDPTS	14127	4	UL-PEN	18/06/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
ATS	16133	576	UL-PEN	18/06/76	FTN	4.5	414	665X	I	665X	I	OPT=2	
/PCL.C./	1391	23											
/SH-10./	16754	131	SI-FORTAN	18/01/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
COMIGs	15185	64	SI-FORTAN	08/11/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
ENTRATZ	15185	8	SI-FORTAN	08/11/76	COMPASS	3	2-414	665X	I	665X	I	OPT=2	
EOF	15171	16	SI-FORTAN	08/11/76	COMPASS	3	2-414	665X	I	665X	I	TEST FOR END OF FILE STATUS.	

Table B2. Loader Map for TPILOT Program (Cont.)

Table B2. Loader Map for TPLOT Program (Cont.)

0350 357015 M3 033599

ESTATE PLANNING

REPELLINE INSTITUTE REQUEST.

Table B1. Sample Output From TPILOT Program

ENCL OF PLG C/S

Table B4. Data Cards for TPLOT Program

CLOUD LORBERTZ(48) JULY 22, 77

3338. 3658. 48.